

A nighttime photograph of the Chain Bridge in Budapest, Hungary, illuminated with warm lights and reflected in the water.

Missions to Habitable Worlds

28-29 October 2015
Budapest, Hungary

International conference
on astrobiology issues related
to next European space missions.
Organized by COST TD 1308 action

Astrobiology in ESA missions

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ESA-ESTEC

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- **The ESA Scientific Programme**
- **Running missions**
- **Missions under development & study**
- **Plans for the future**

22 Member States and growing



ESA has 22 Member States: Austria, Belgium, Czech Republic, Denmark, Estonia, Finland, France, Germany, Greece, Hungary, Ireland, Italy, Luxembourg, The Netherlands, Norway, Poland, Portugal, Romania, Spain, Sweden, Switzerland and the United Kingdom.

Canada also sits on the Council and takes part in some projects under a Cooperation Agreement. Other EU states also have Cooperation Agreements with ESA, such as Bulgaria, Cyprus, Lithuania and Malta. Latvia, Slovenia and Slovakia are participating in the Plan for European Cooperating States (PECS).



SCIENCE & ROBOTIC EXPLORATION



Science-driven

both long-term science planning and mission calls are bottom-up processes, relying on broad community input and peer review.

Mandatory

all member states contribute pro-rata to GNP providing budget stability, allowing long-term planning of its scientific goals and being the backbone of the Agency.

European Space Agency

→ ESA'S FLEET IN THE SOLAR SYSTEM

The Solar System is a natural laboratory that allows scientists to explore the nature of the Sun, the planets and their moons, as well as comets and asteroids. ESA's missions have transformed our view of the celestial neighbourhood, visiting Mars, Venus, and Saturn's moon Titan, and providing new insight into how the Sun interacts with Earth and its neighbours. The Solar System is the result of 4.6 billion years of formation and evolution. Studying how it appears now allows us to unlock the mysteries of its past and to predict how the various bodies will change in the future.

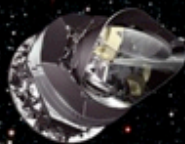
→ ESA'S FLEET ACROSS THE SPECTRUM



Thanks to cutting edge technology, astronomy is unveiling a new world around us. With ESA's fleet of spacecraft, we can explore the full spectrum of light and probe the fundamental physics that underlies our entire Universe. From cool and dusty star formation revealed only at infrared wavelengths, to hot and violent high-energy phenomena, ESA missions are charting our cosmos and even looking back to the dawn of time to discover more about our place in space.

planck

Looking back
at the dawn of time



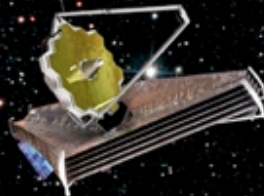
herschel

Unveiling the cool
and dusty Universe



jwst

Observing the first light



cheops

Sizing and first characterisation
of exoplanets



euclid

Exploring the dark Universe



gaia

Surveying a billion stars



hst

Expanding the frontiers
of the visible Universe



xmm-newton

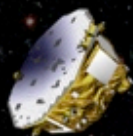
Seeing deeply into the hot
and violent Universe



lisa

pathfinder

Testing the technology
for gravitational
wave detection



integral

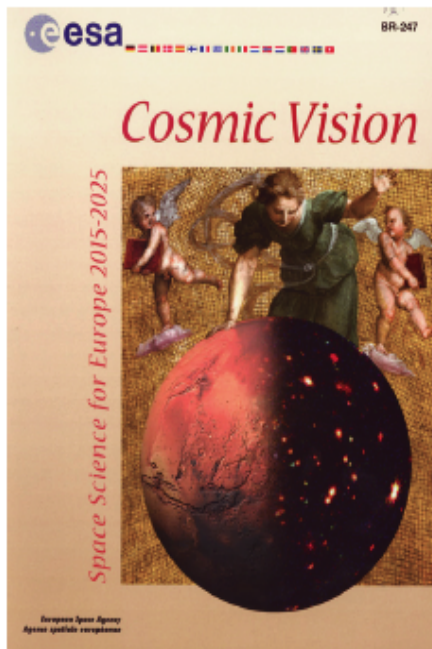
Seeking out the extremes
of the Universe



ESA's long-term scientific programme is based on a vision.

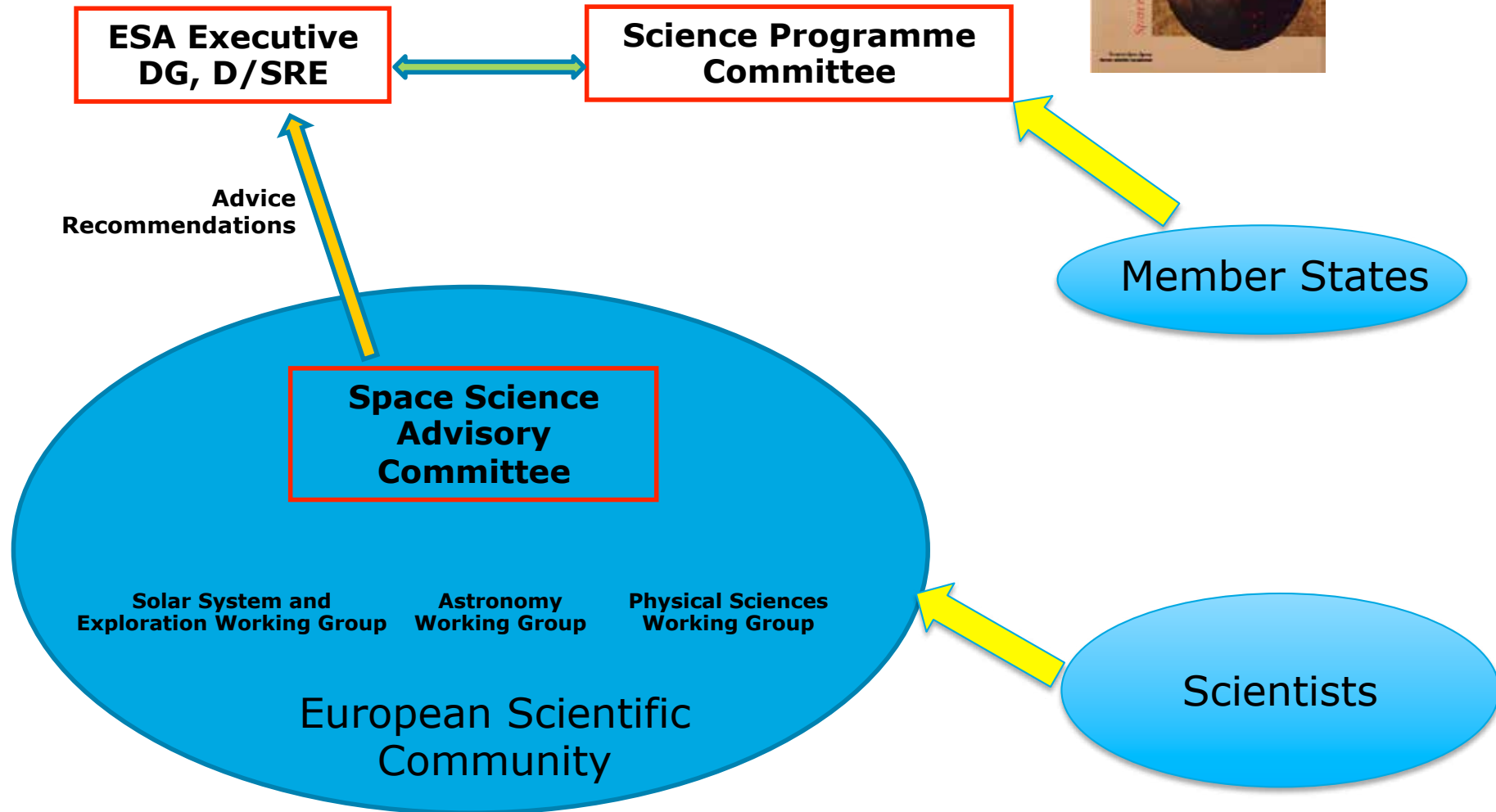
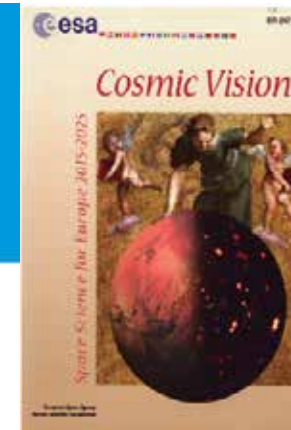
The 'Cosmic Vision' looks for answers to mankind's fundamental questions:

- What are the conditions for planet formation and the emergence of life?
- How does the Solar System work?
- What are the fundamental physical laws of the Universe?
- How did the Universe originate and what is it made of?



COSMIC VISION

A bottom-up approach



Science Programme building blocks “Large” (Ariane 5-class) missions



1. High innovation content
2. European flagships
3. 3 per 20 years



Science Programme building blocks “Medium” (Soyuz-class) missions



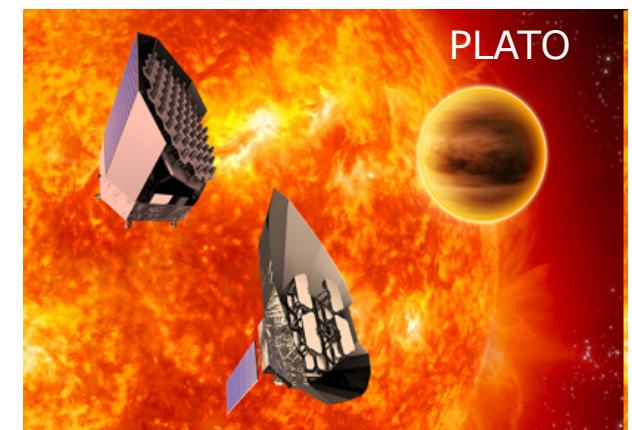
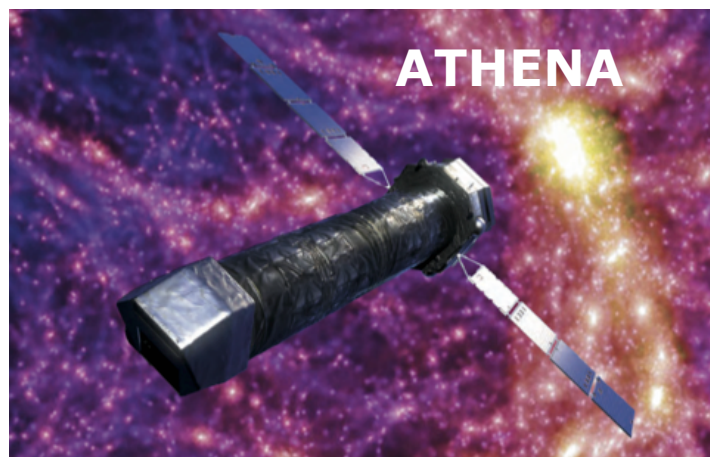
1. Makes use of current cutting-edge technology
2. Programme workhorse
3. 3-4 per 10 years



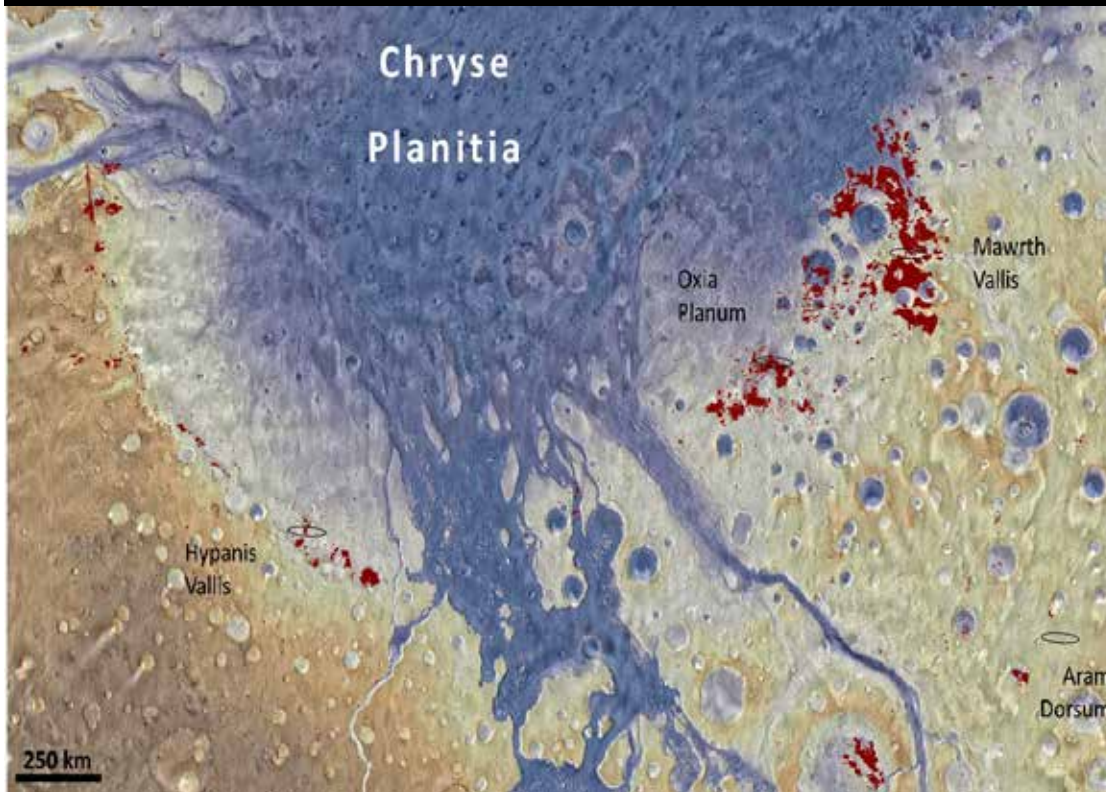
COSMIC VISION



- Selection of Solar Orbiter as M1 and Euclid as M2 in 2011
- Selection of JUICE as L1 in 2012
- Selection of CHEOPS as S1 in 2012
- Selection of PLATO in early 2014 as M3
- Selection of ATHENA in June 2014 as L2



OMEGA: Mapping hydrated minerals and sulfates

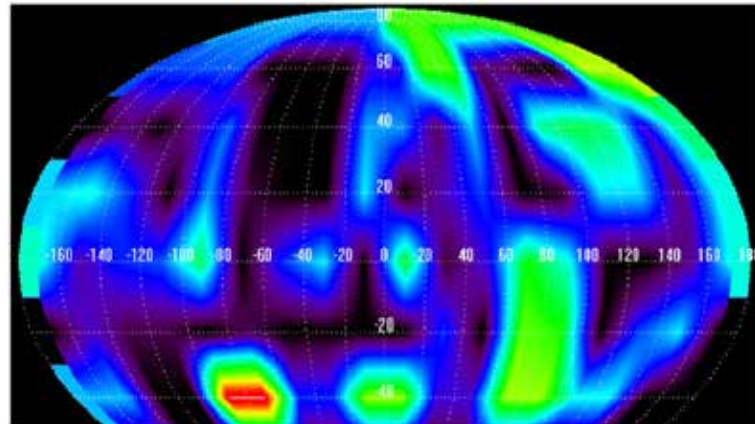


➤ Reconstruction of the Martian paleo-climate

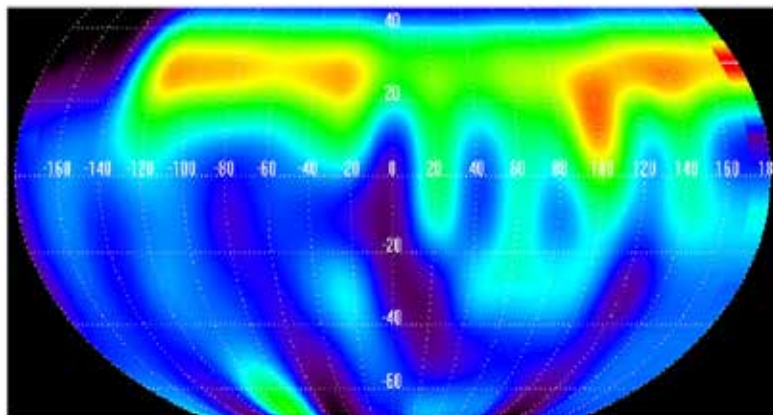
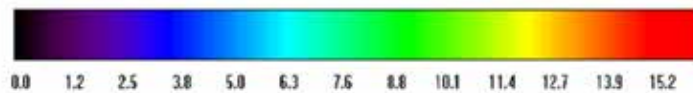
- Clays formed in Noachian (4.2-3.7 Gy) → wet and warm climate
- Sulfates formed in Hesperian (3.7-3.0 Gy) → dry and more acid climate

PFS:

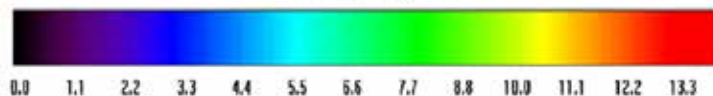
Seasonal variations of methane



CH₄ column density * e-15



CH₄ column density * e-15



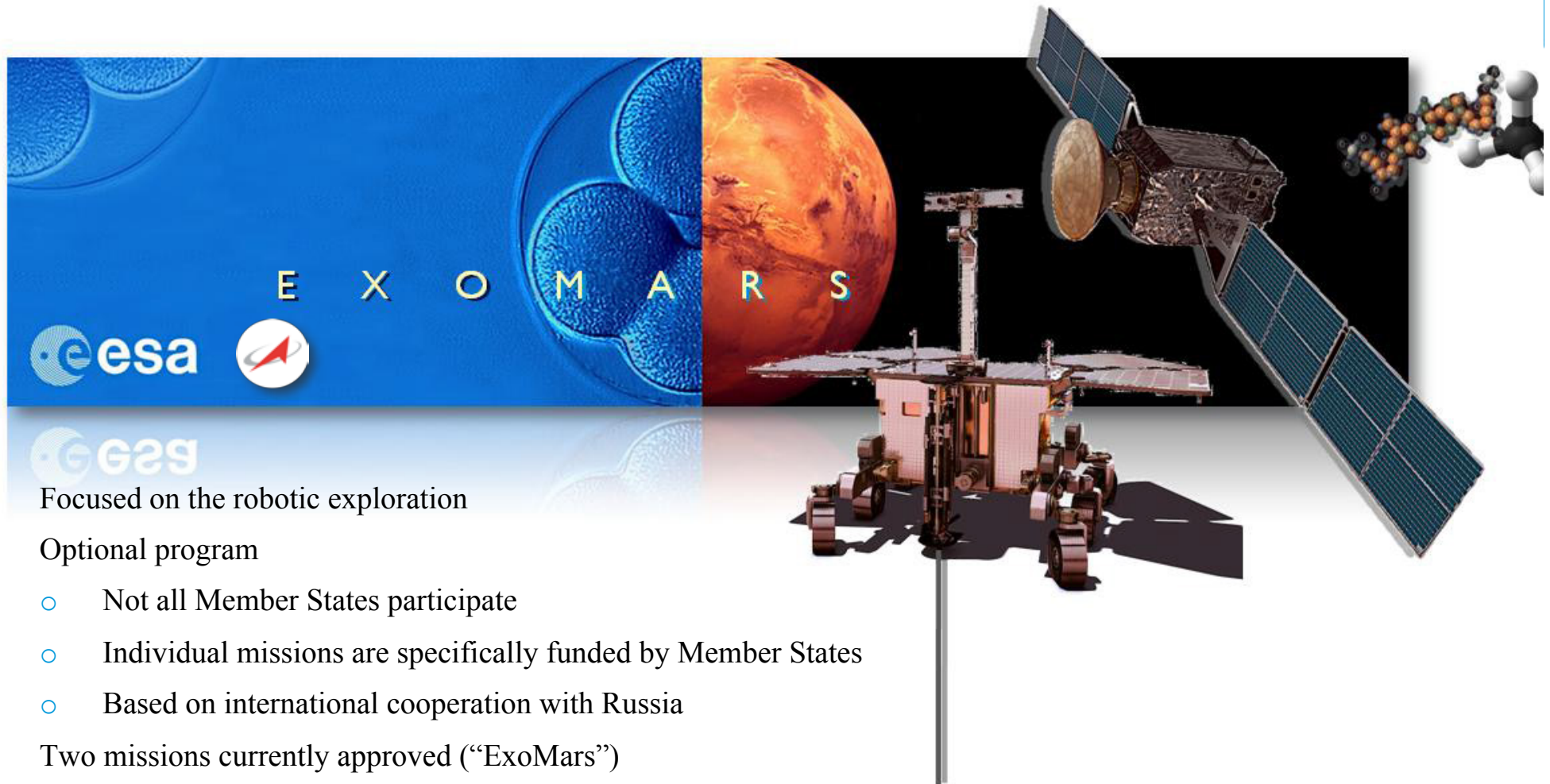
Northern summer

- **Methane mixing ratio 0-45 ppb**
- **Non-homogeneous distribution suggesting sources and sinks**
- **Increase of methane column density occurs in the winter hemisphere**
- **Observations of methane in Gale crater (Curiosity)**

Northern winter

Formisano et al., 2004; Geminale et al., 2008, 2011

The (optional) European Robotic Exploration Program



- Focused on the robotic exploration
- Optional program
 - Not all Member States participate
 - Individual missions are specifically funded by Member States
 - Based on international cooperation with Russia
- Two missions currently approved (“ExoMars”)
 - Trace gas orbiter (TGO) and Entry, Descent, and Landing Demonstrator Module (EDM) (2016)
 - Exo-biology rover with Pasteur P/L plus Surface Platform (2018)
- Long-term goal is Mars Sample Return

First landing on a world in the outer Solar System

In 2005, ESA's Huygens probe made the most distant landing ever, on Titan, the largest moon of Saturn (about 1427 million km from the Sun).



EXO BIOLOGY

The study of Life and related structures and processes in the Universe

TITAN's Geofluid

a prebiotic reactor at a planetary scale

Titan's Chemistry & Exobiology

Organic Chemistry in the atmospheric **gas phase**

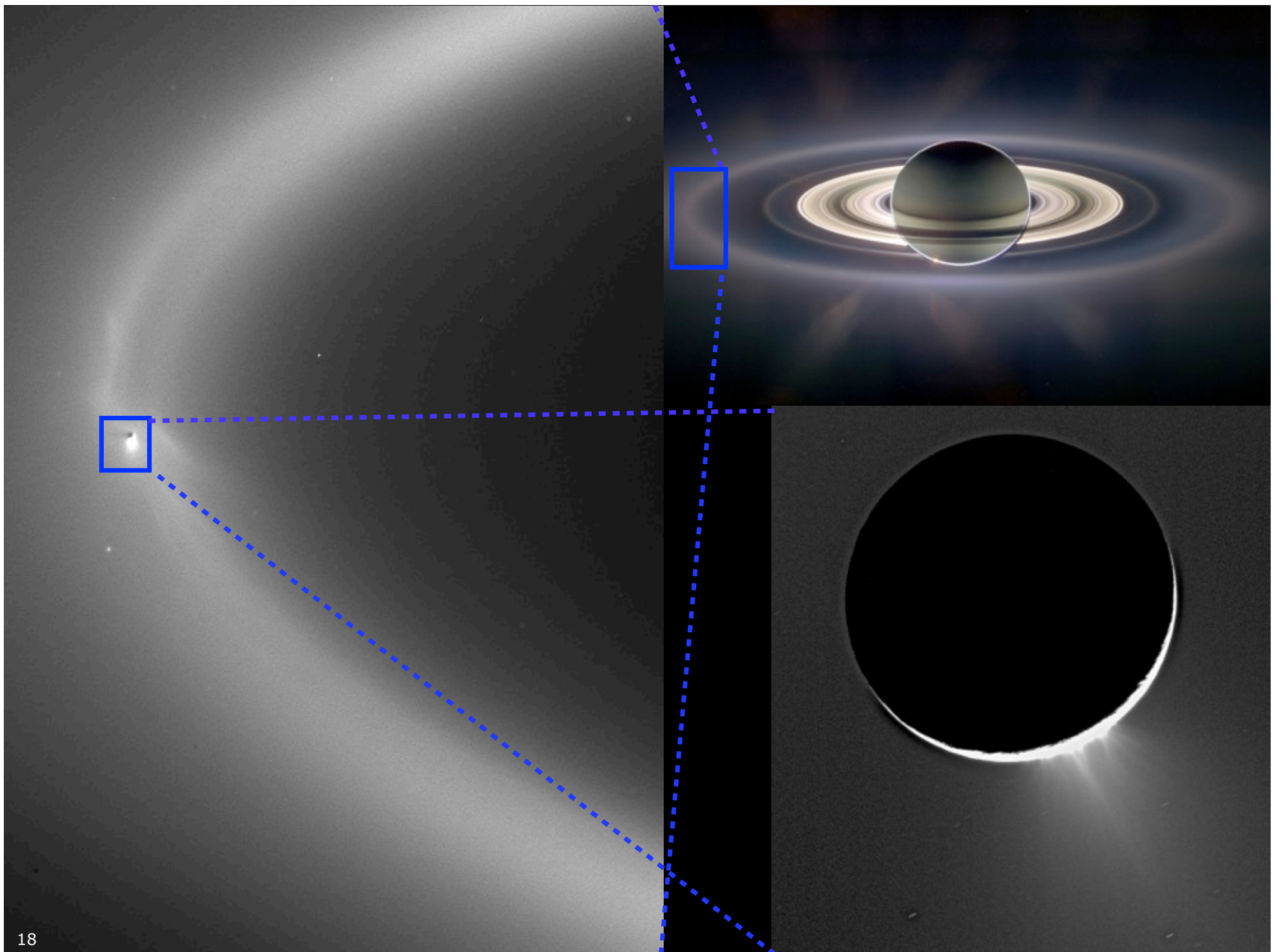
Organic Chemistry in the atmospheric **aerosols**

Organic Chemistry on the **surface** & in the **sub-surface**

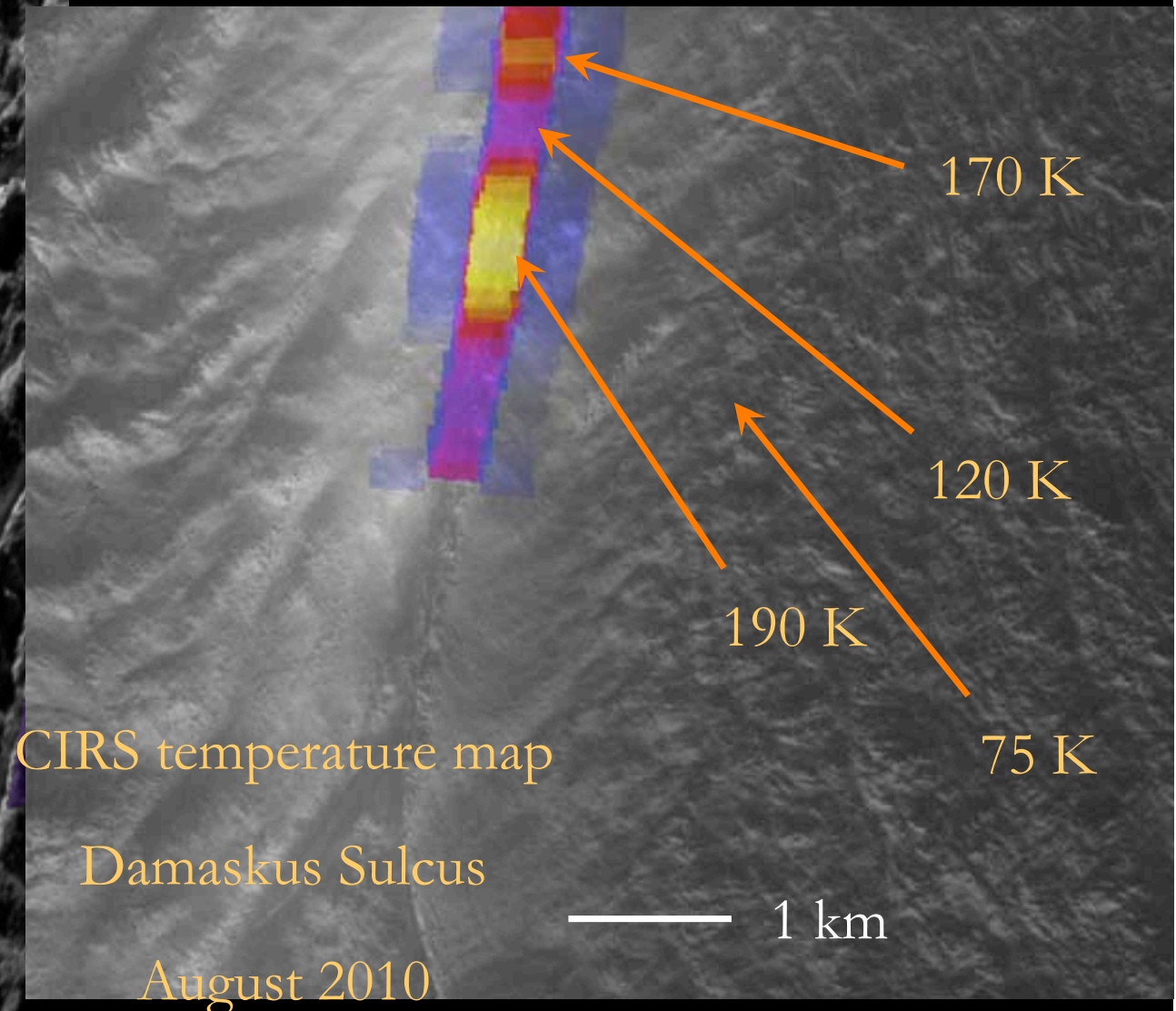
Physical & Chemical Couplings



**Enceladus astrobiological assets:
An icy-crust, a sub-surface ocean, a rocky core, and
heat**



Heat: generated by tidally driven frictions, due to Enceladus small orbit eccentricity around Saturn

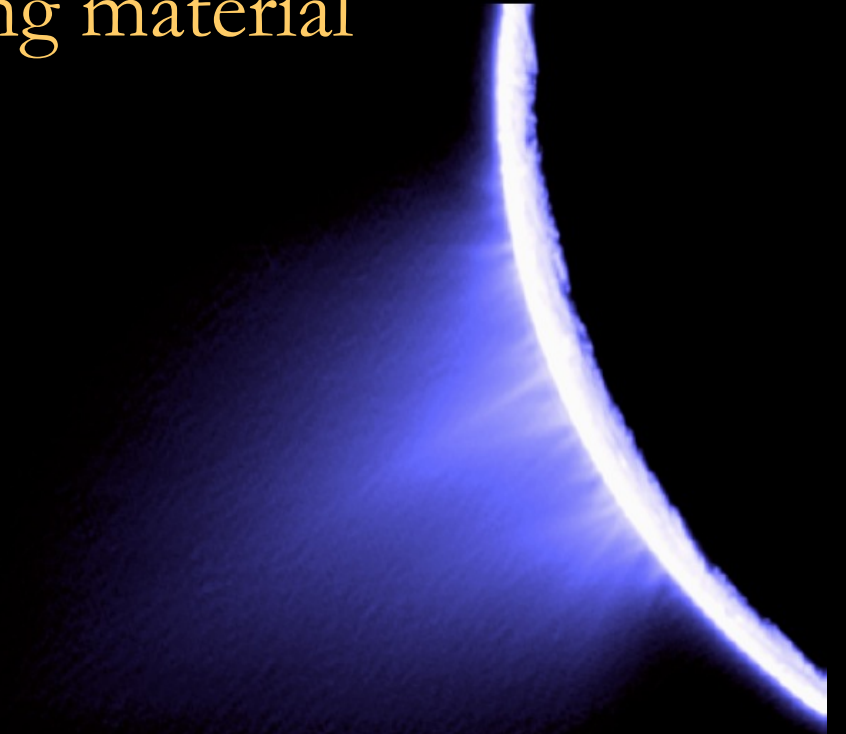


Composition of escaping material

- Gas phase:

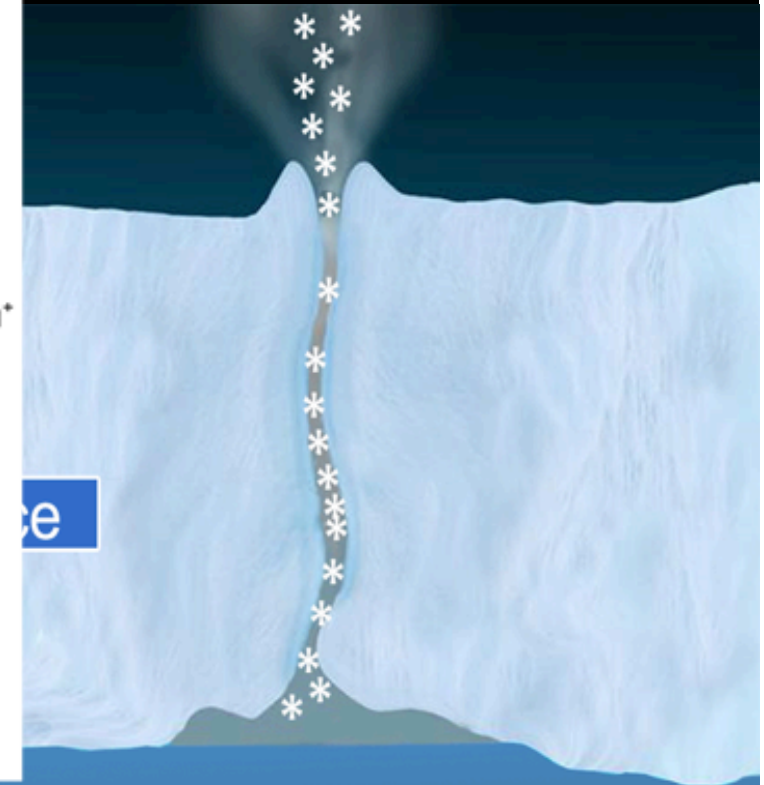
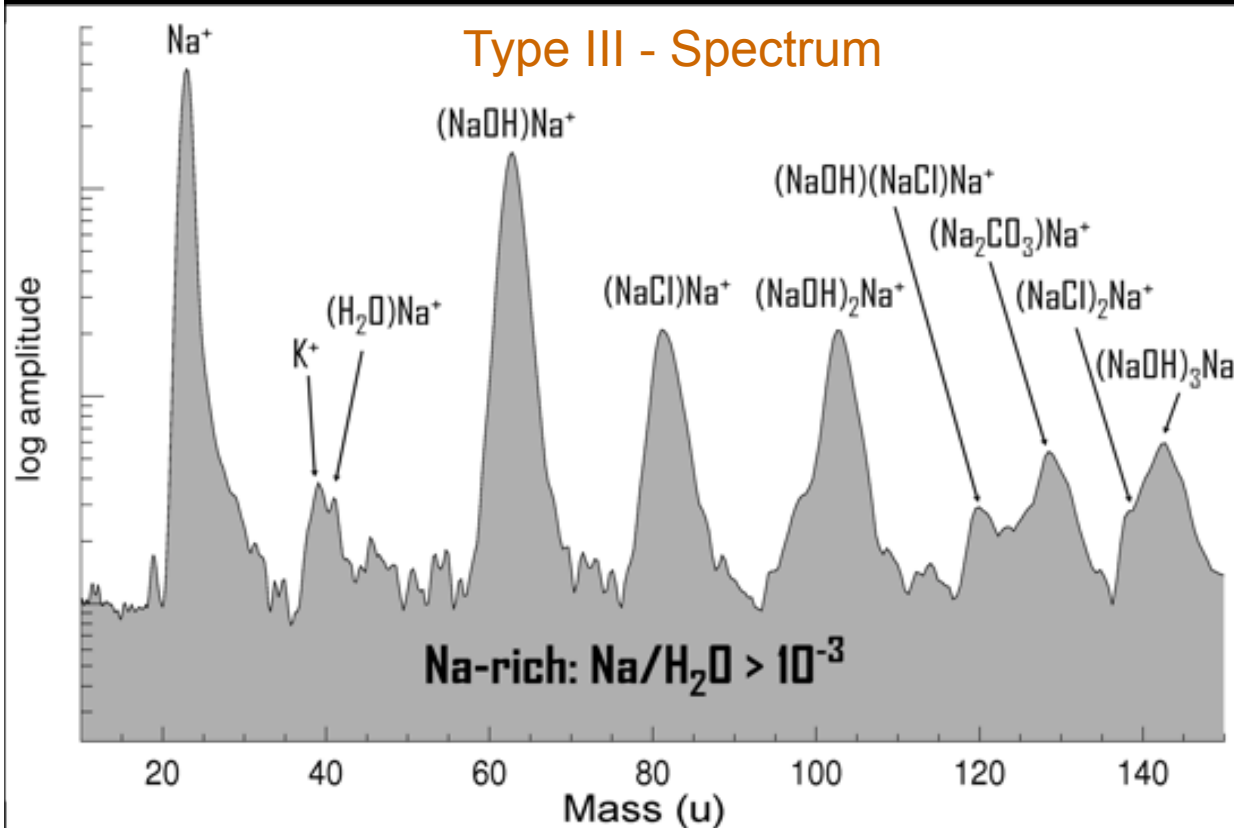
- H_2O $> 92 \%$
- CO_2 $\sim 0.5 - 5\%$
- volatile organics $\sim 1 - 3 \%$
- NH_3 $\sim 1\%$
- CO $< 3\%$
- N_2 $< 0.5 \%$
- Na $< 0.0001 \%$

Waite et al, Science, 2006



Composition of escaping material: solid phase

Postberg et al. Nature, 2011



⇒ 98% ice, 1 - 2% salts

⇒ NaCl , NaHCO_3 / Na_2CO_3 , KCl

→ saltwater with alkaline pH (8 – 9)

→ The water must be in contact with rocky core

Water

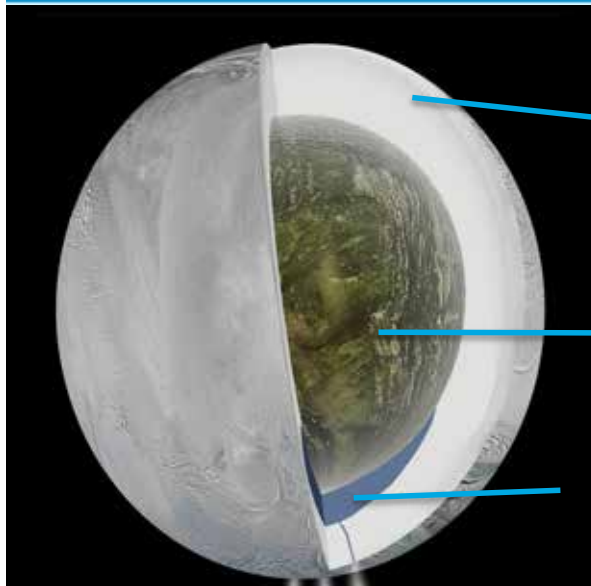
Na^+ Cl^- HCO_3^- CO_3^{2-} K^+

Rock

NASA / JPL

Characterizing Enceladus sub-surface ocean

Ocean extent and hydrothermal activity

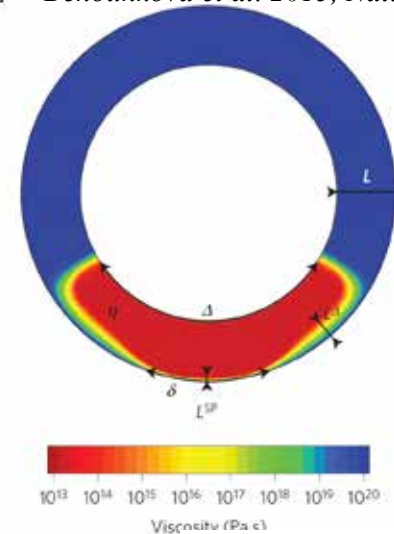


Outer ice shell
50 km thick

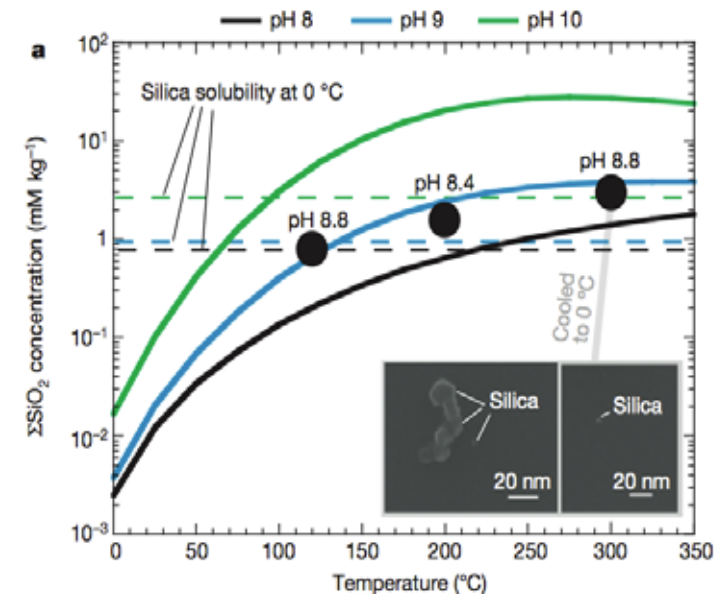
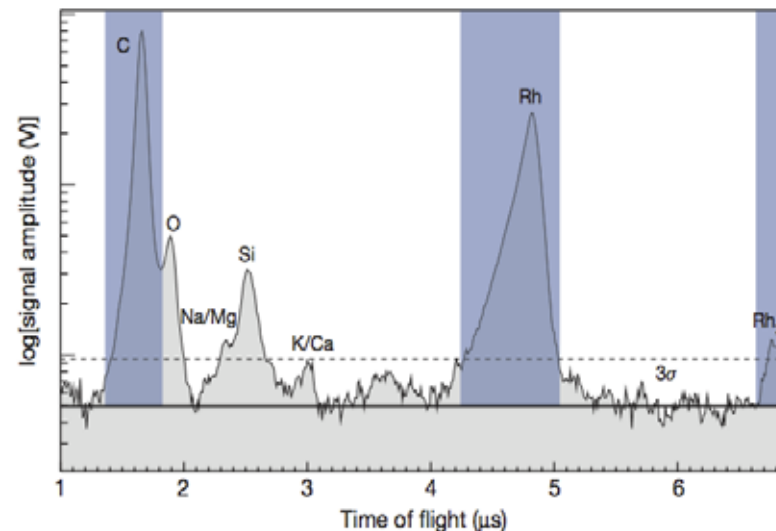
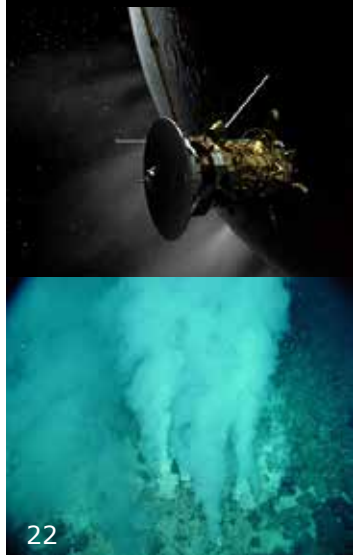
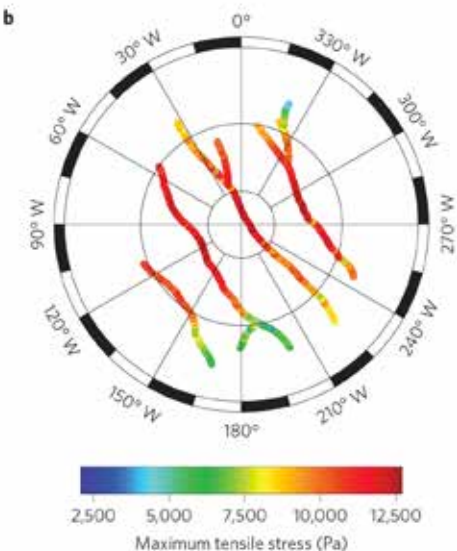
Silicate core
density 2.4
radius 200 km

Regional ocean (??)
8-10 km deep, extending
to 50° S

a Behoukova et al. 2015, Nature



b

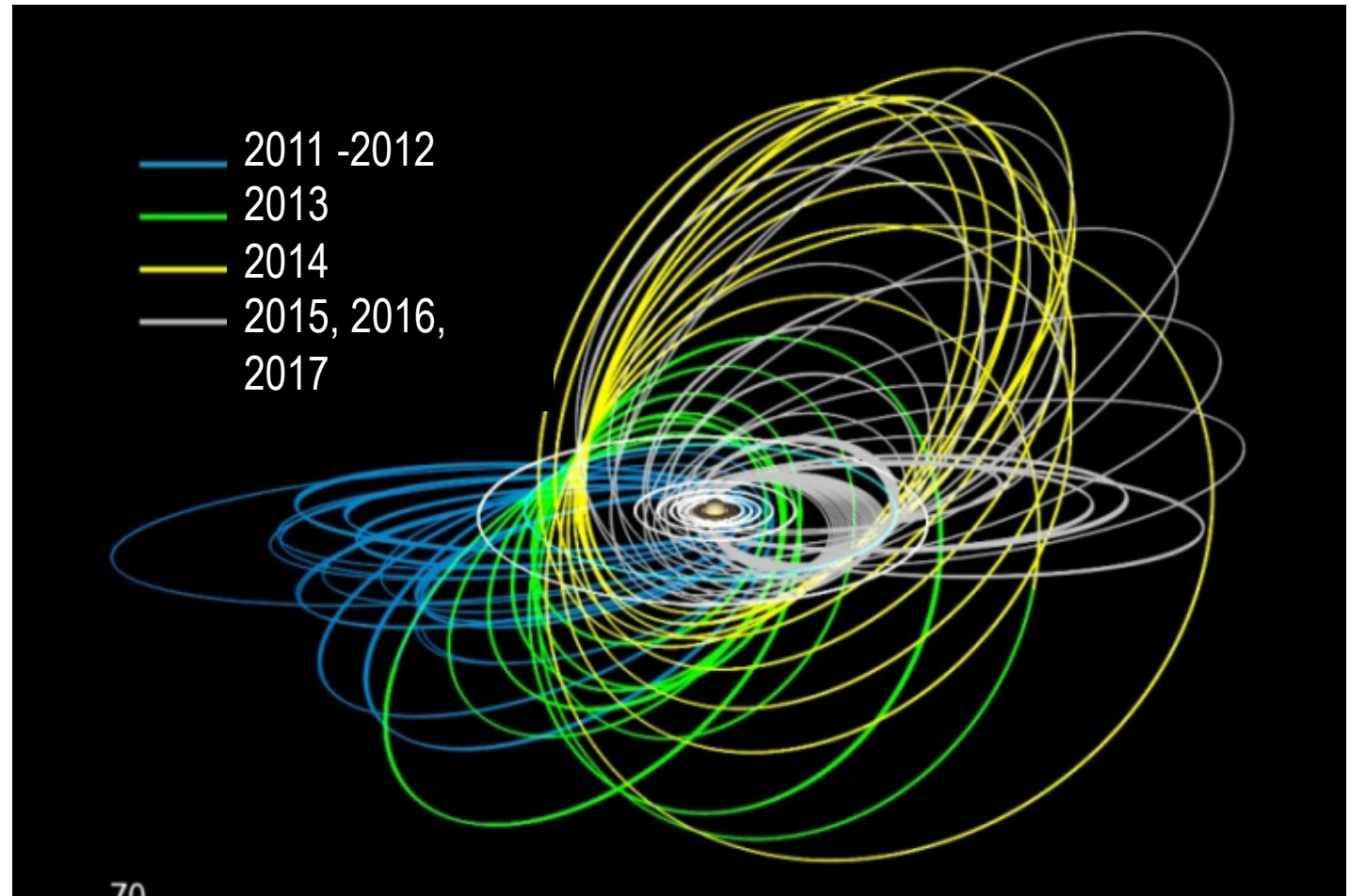


Hsu et al. – Cassini-CDA team, 2015, Nature

Cassini Mission Overview: Trajectory



- ❑ Cassini is a ‘Touring’ mission
- ❑ Large variety of orbits, within and outside the equatorial plane
- ❑ Inclination steps changes reached through Titan targeted flybys [scientific purposes/fuel efficient navigation]



JUICE: JUpiter Icy moons Explorer



European Space Agency
Agence spatiale européenne

Emergence of habitable worlds around gas giants

Jupiter system as an archetype for gas giants

Callisto:

remnant of the early solar system

- Icy shell, ocean
- Geology, surface composition
- Past activity

Europa: recently active zones

- Surface non-water-ice material
- Search for liquid water
- Recent activity

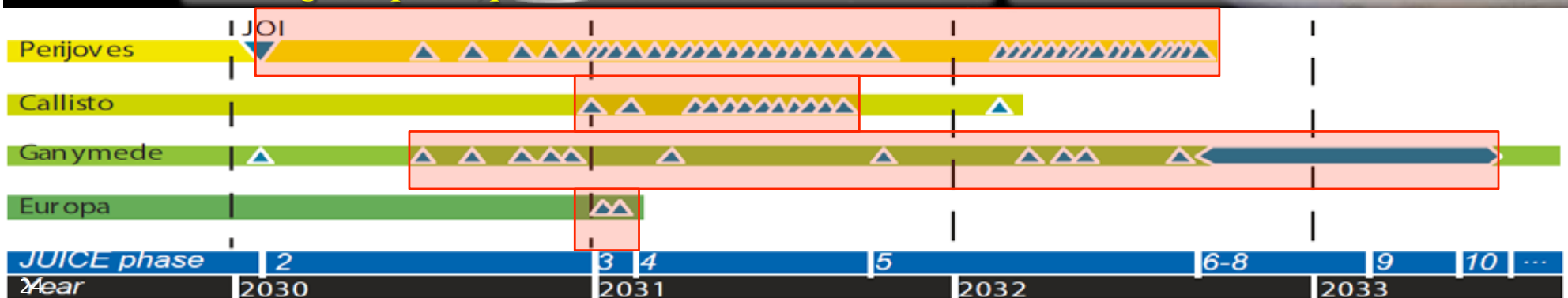
Ganymede:

planetary object and potential habitat

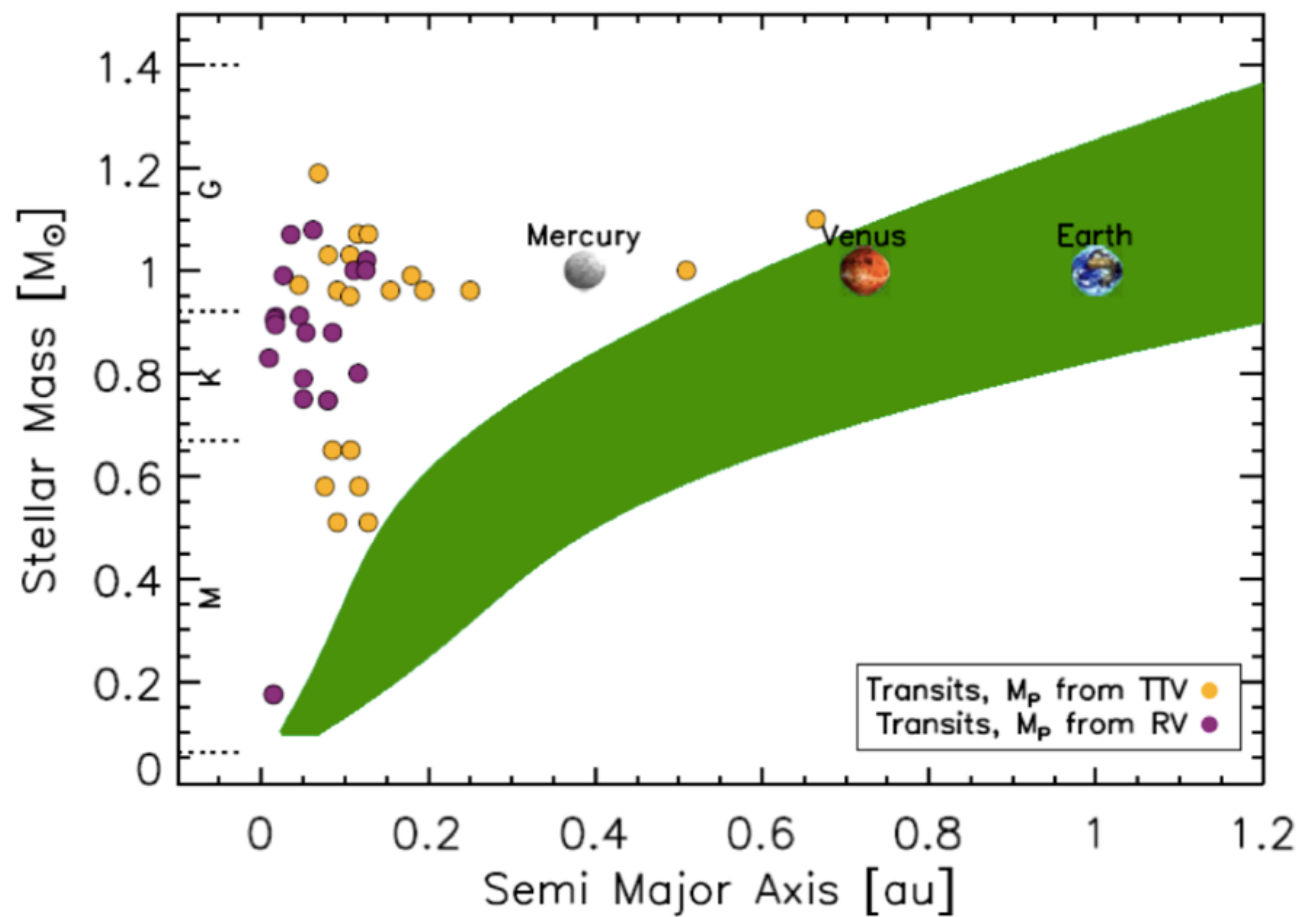
- Sub-surface, ice shell, ocean, interiors
- Geology, surface composition
- Atmosphere, ionosphere
- Magnetosphere, plasma environment

Jupiter System:

- Atmospheric structure, chemistry and dynamics
- Magnetosphere as fast rotator and giant accelerator
- Moons as plasma sources and sinks
- Couplings and interactions

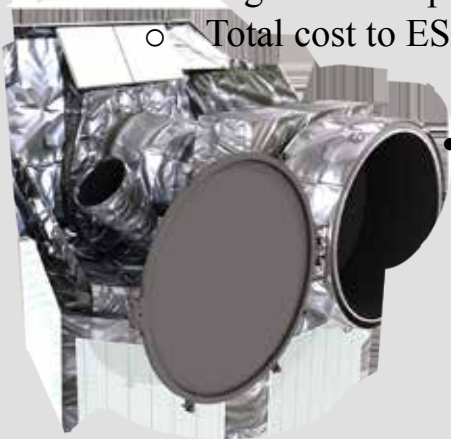


Exo planets



CHEOPS: an S-class mission

- First S (small) - class mission in the ESA Science Programme
- Implemented in partnership with Switzerland, with a consortium comprising a large number of ESA member states (Austria, Belgium, France, Germany, Hungary, Italy, Portugal, Spain, Sweden, UK)
- Boundary conditions:
 - High TRL for platform and payload → development time < 3.5 – 4 yrs
 - Total cost to ESA limited to about 0.1 yearly budget



• Selected December 2012, adopted June 2014, launch-ready end 2017 – fast!



CHEOPS science

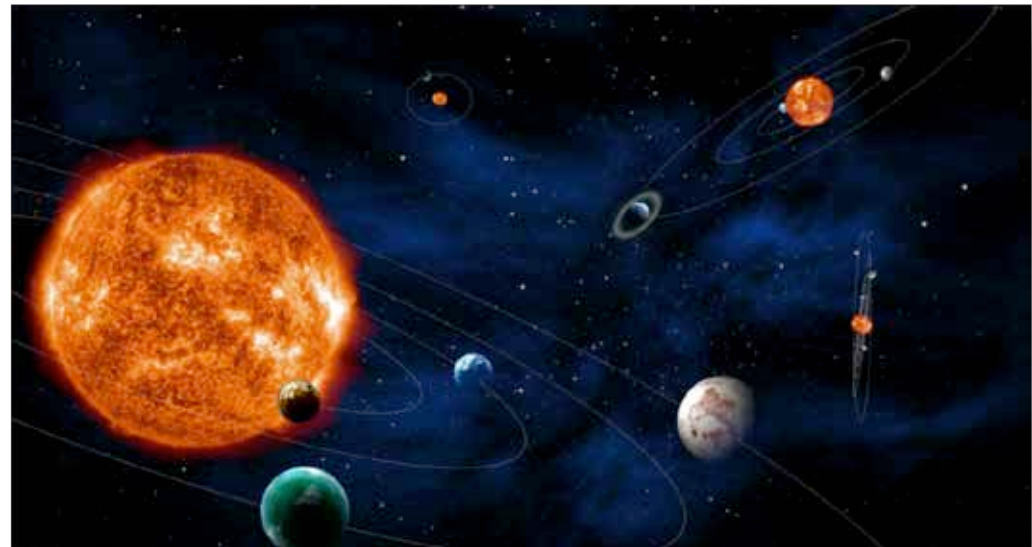
- CHEOPS (CHAracterising ExOPlanet Satellite) – a mission dedicated to the search for exoplanet transits of local, bright stars already known to host exoplanets:
 - Detection and characterisation of transiting exoplanets with masses $< 30 M_{\text{earth}}$ through precision, wideband transit photometry
 - Follow-up, pointed observations:
 - Known exoplanets \rightarrow Follow-up mission \rightarrow Pointed observations, know where and when to point
 - Bright host stars ($V < 12$) \rightarrow precise mass measurements available /feasible, detailed knowledge of the star
 - First-step characterisation of super-Earths and Neptunes: precision masses+radii \rightarrow measurement of bulk density
 - Insight into physics and formation of planets
 - Identification of planets with atmospheres
 - Constraints on planet migration
 - Identification of “golden targets” for spectroscopic characterisation
 - Probing atmospheres of hot-Jupiters using phase curve measurements
 - Study of physical mechanisms and efficiency of energy transport

PLATO

Scientific Objectives

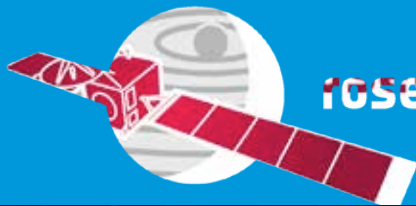


- Detect and characterise Earth-like planets (1-10 Earth masses, 1-2 Earth radii) in the habitable zone of bright solar like stars:
 - radii (down to 2% accuracy, photometric transit method)
 - masses ($\sim 10\%$ accuracy, from radial velocity follow-up at ground-based telescopes)
 - mean densities
 - ages ($\sim 10\%$ accuracy, astero-seismology analysis)
 - host stars knowledge
- Detect and characterise thousands of rocky, icy and giant planets, the architecture of their planetary system and their host star
- Advance stellar science



- Planets in the habitable zone of solar like stars (F5V-K7V)
 - Enough precision in radius and mass to distinguish terrestrial planets from mini-gas planets
=> Identify prime candidates for potentially habitable worlds
 - Enough precision in age to study Earth-like planets at different epochs
- Well characterised planets around the brightest stars for atmospheric spectroscopic observations with other facilities
- Host stars studied with unprecedented accuracy-> Influence of stellar age, stellar type and stellar activity in habitability
- Populations of planets in circumbinary systems and around late type stars
- Exo-moons and their environment

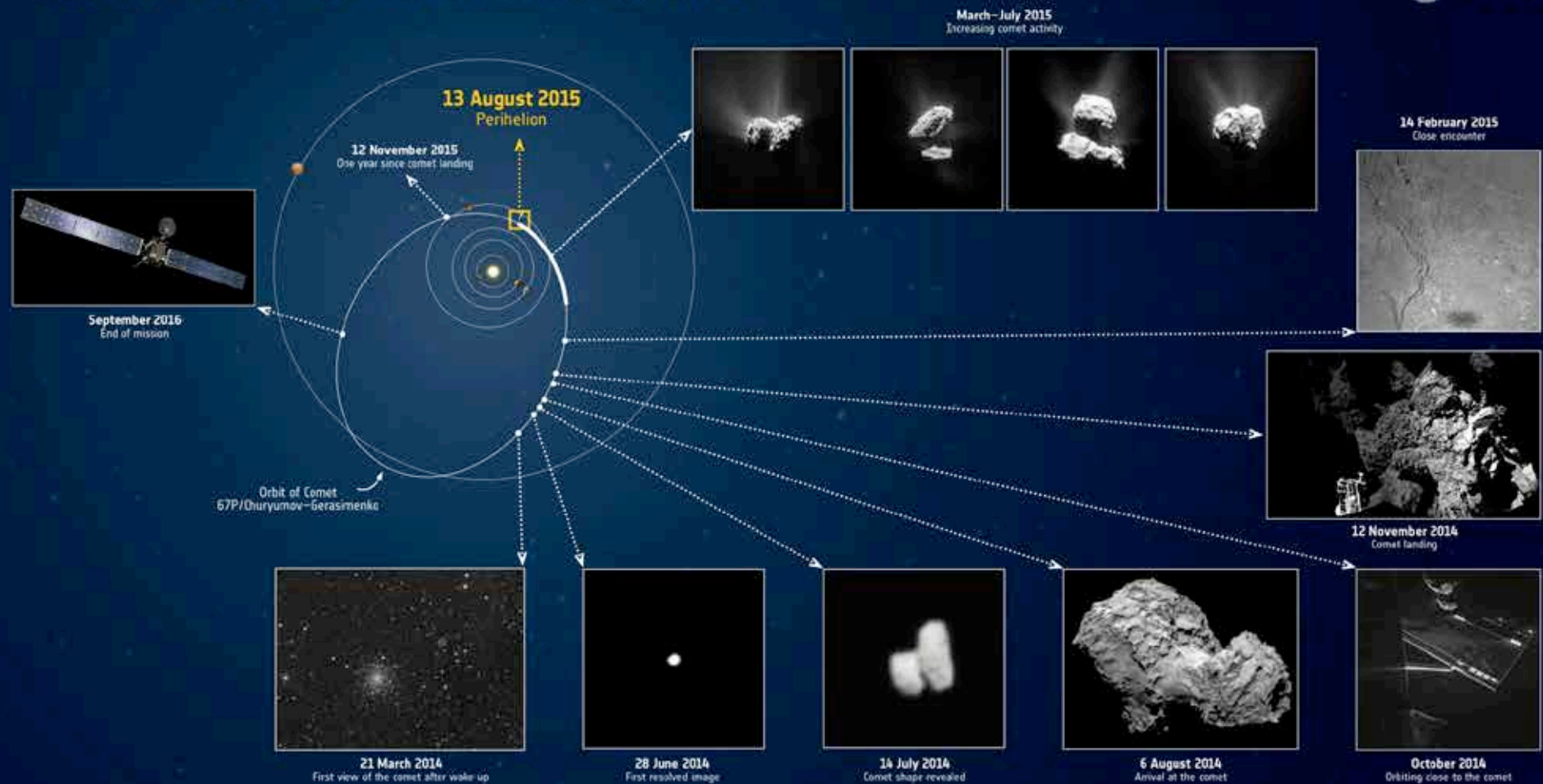




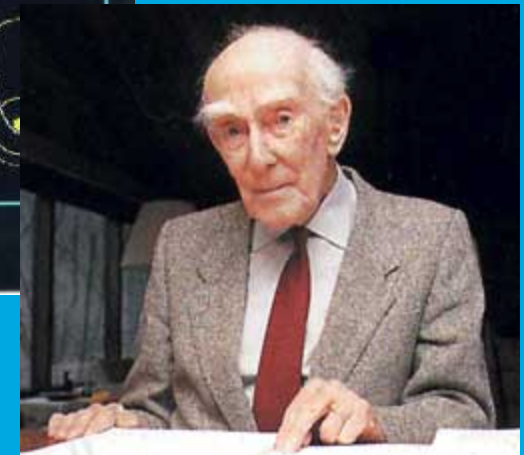
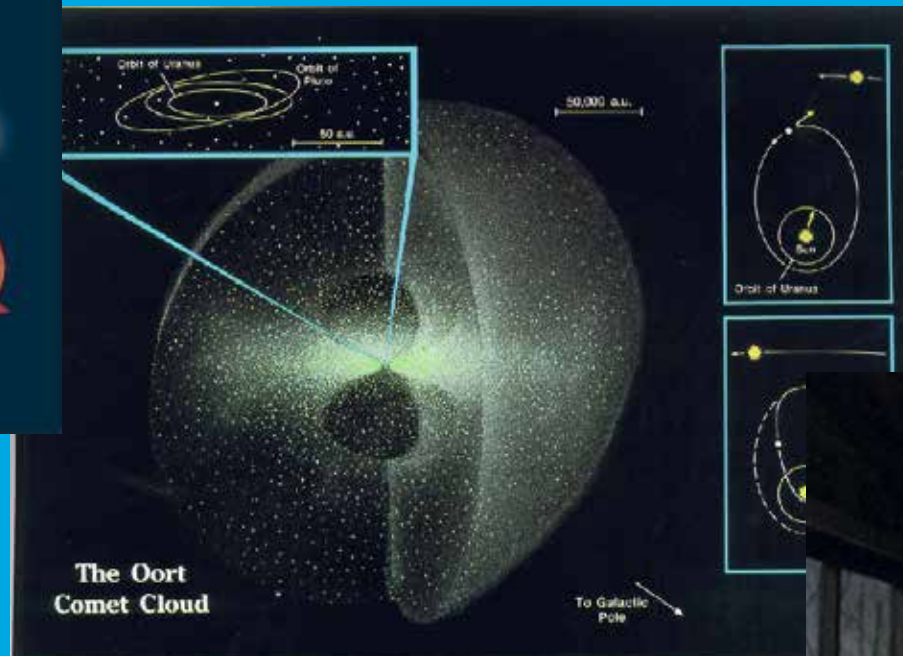
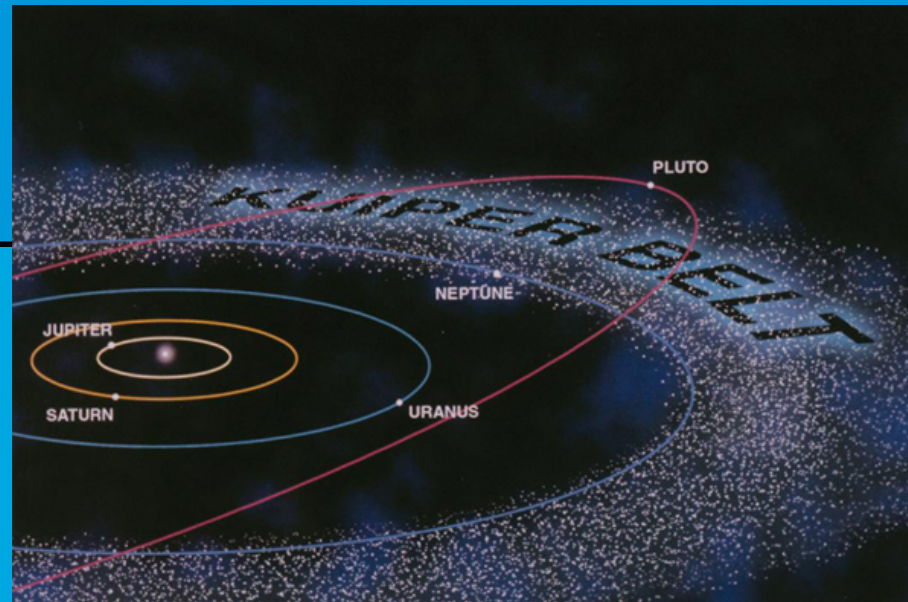
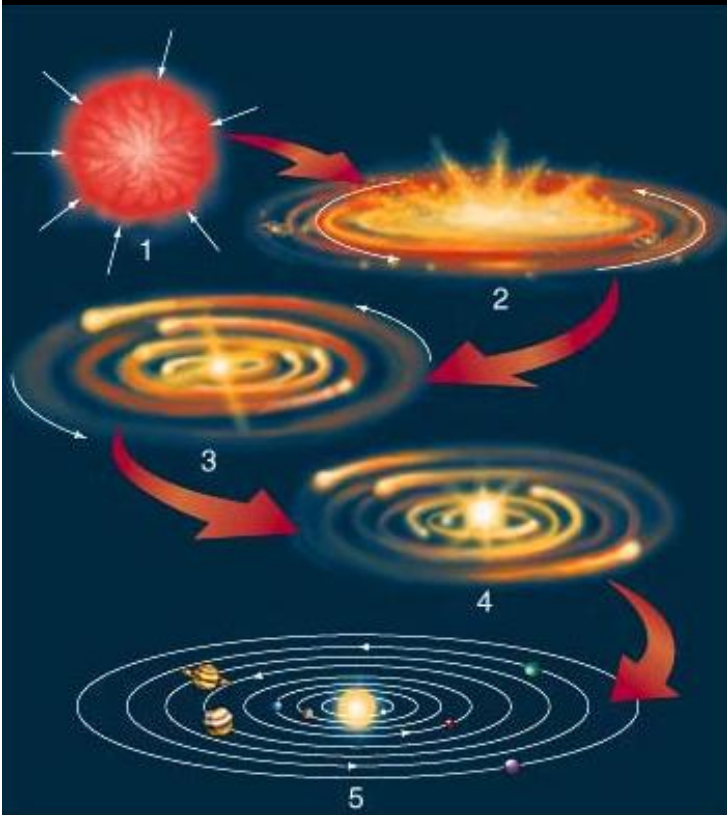
rosetta The ESA Rosetta Mission



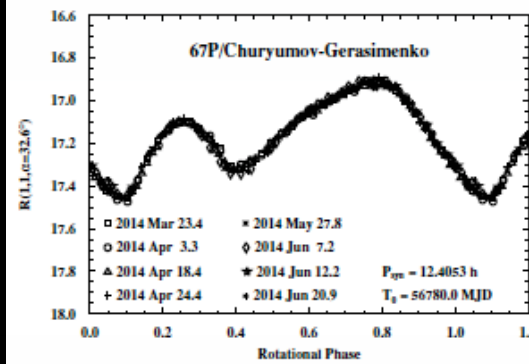
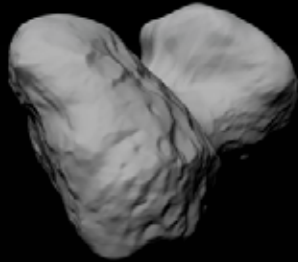
→ ROSETTA: LIVING WITH A COMET



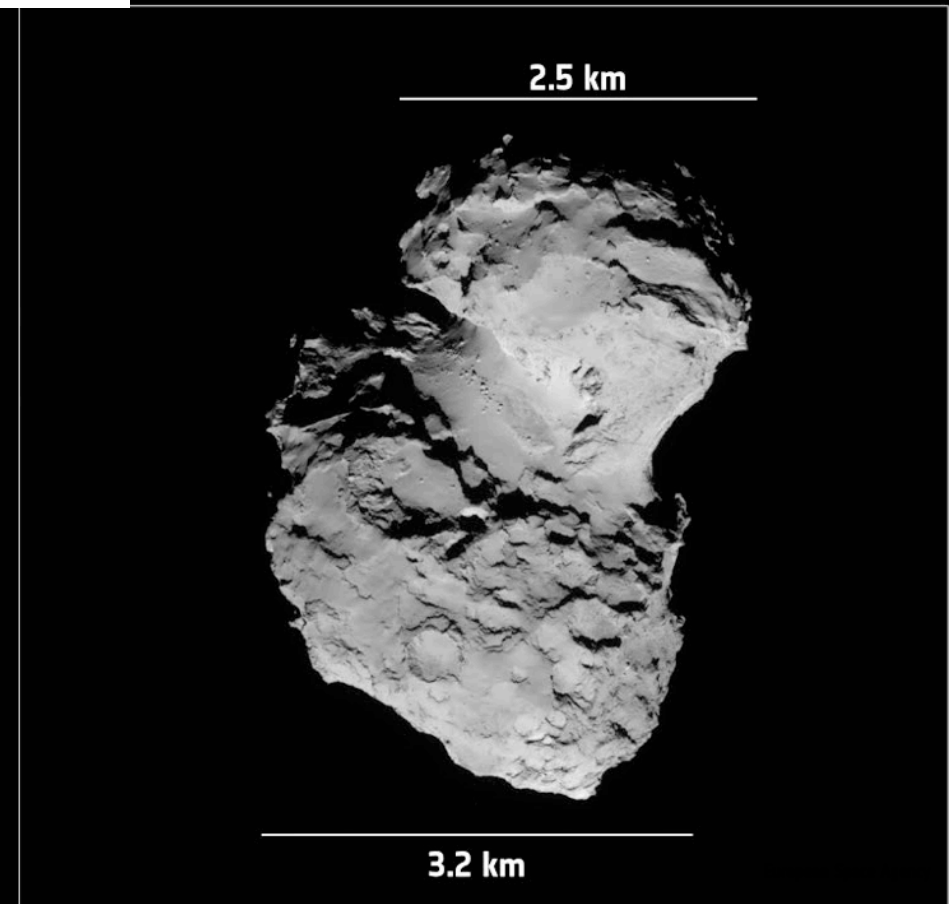
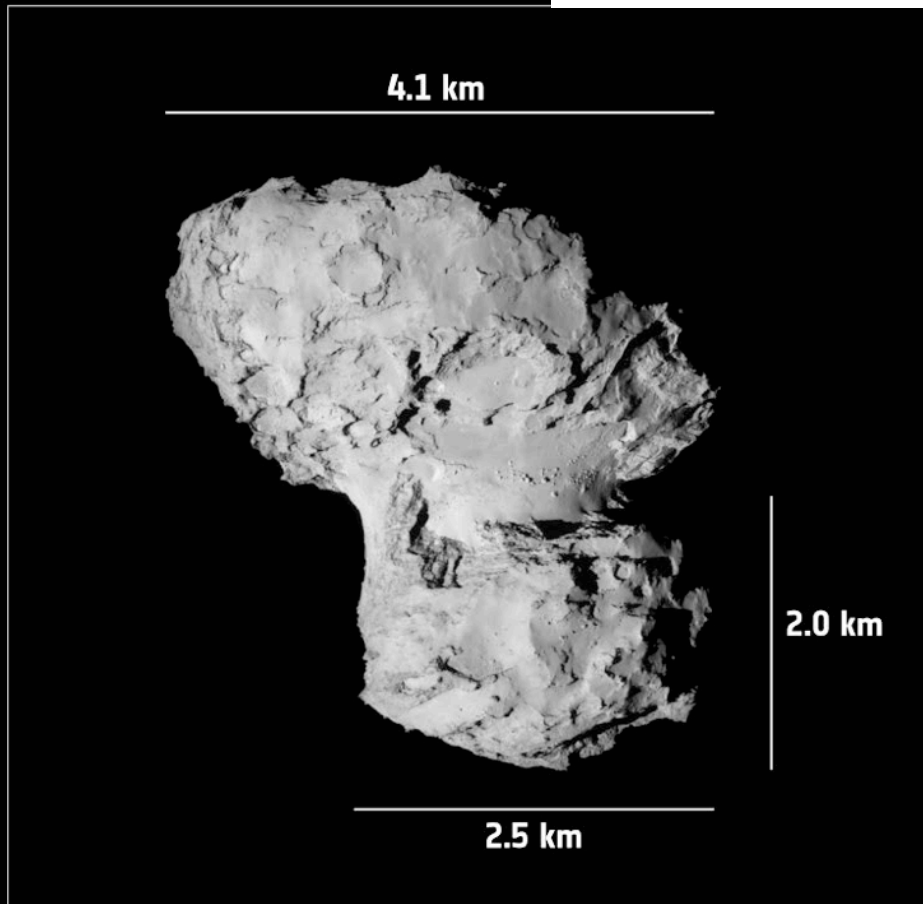
Why comets

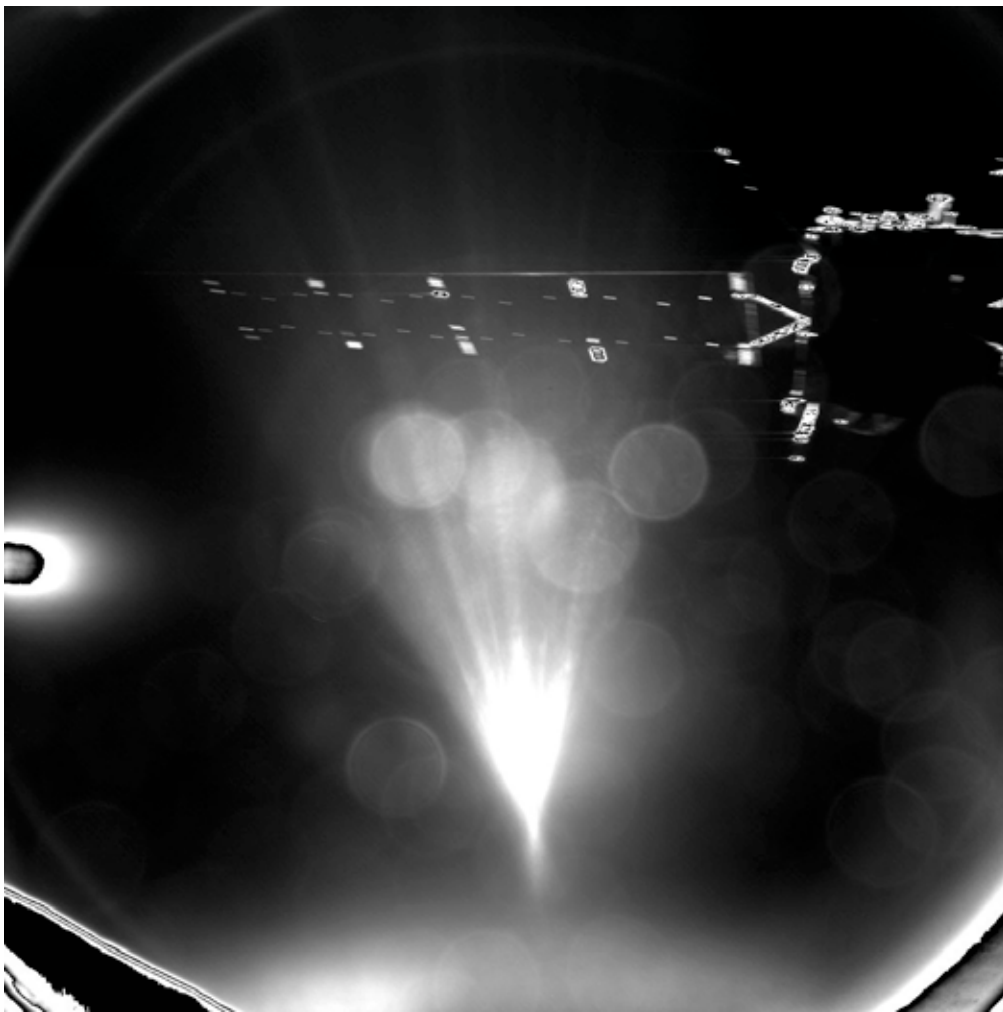


Target: 67P/Churyumov-Gerasimenko

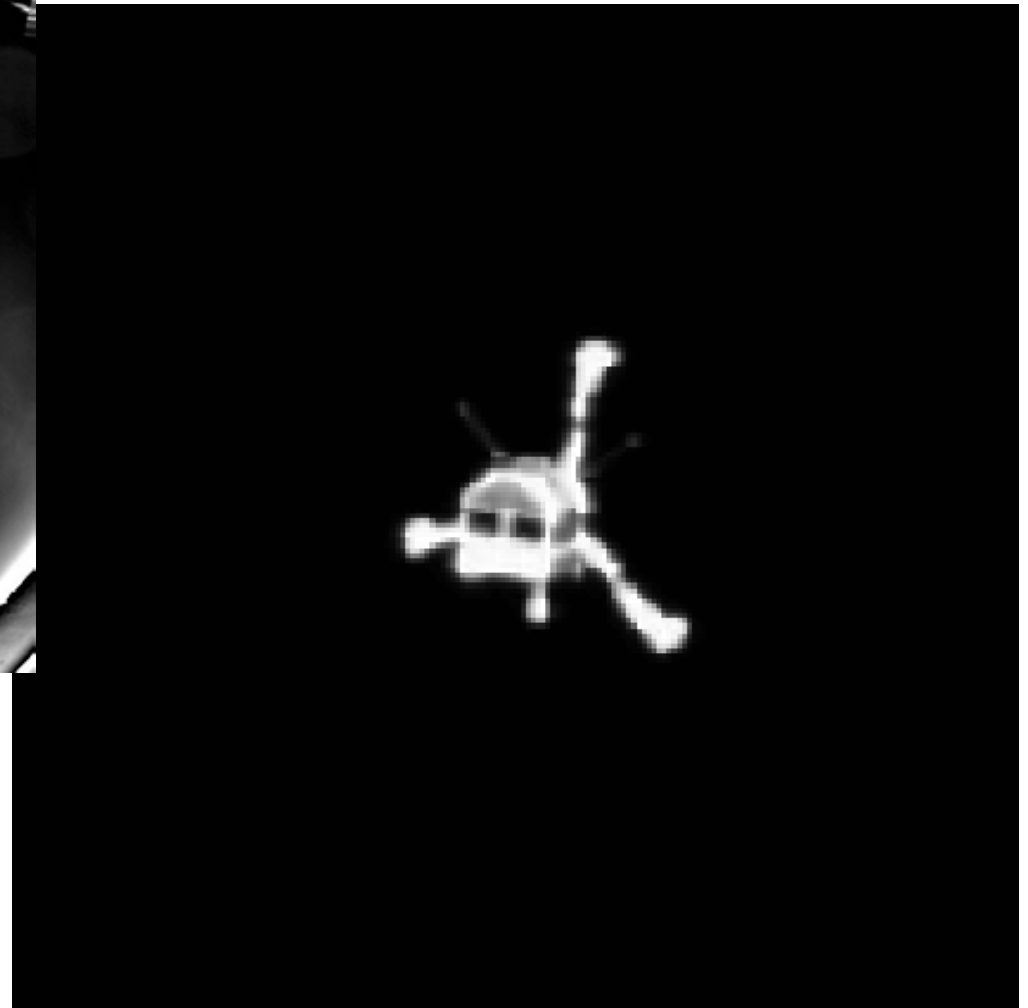


Mass $\sim 10^{13} \text{ kg}$
Volume $\sim 25 \text{ km}^3$
Density $\sim 0.4 \text{ g/cm}^3$
Rotation $\sim 12.4 \text{ hour}$





ESA/Rosetta/Philae/CIVA



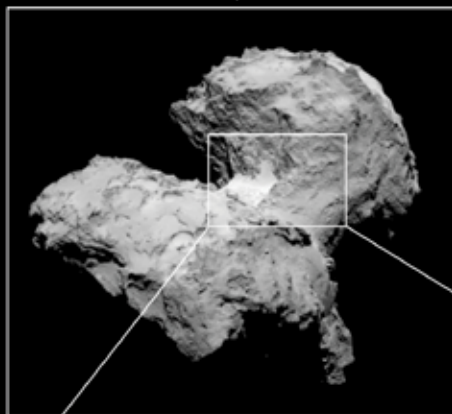
ESA/Rosetta/MPS for OSIRIS Team MPS/UPD/LAM/IAA/SSO/INTA/UPM/DASP/IDA

European Space Agency

→ THE CYCLE OF WATER ICE AT COMET 67P/CHURYUMOV–GERASIMENKO



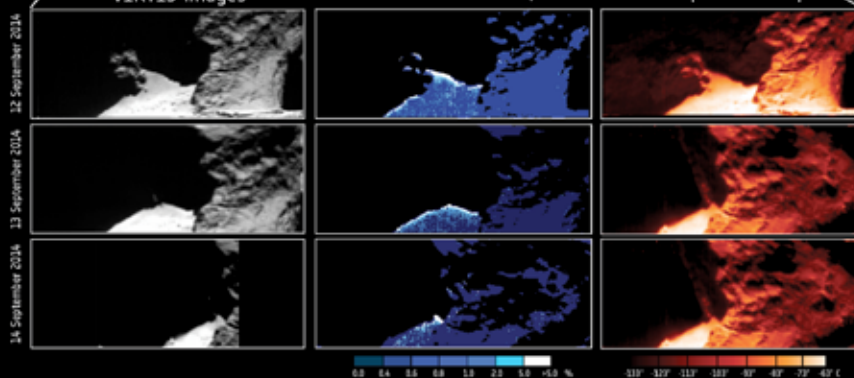
Comet on 2 September 2014



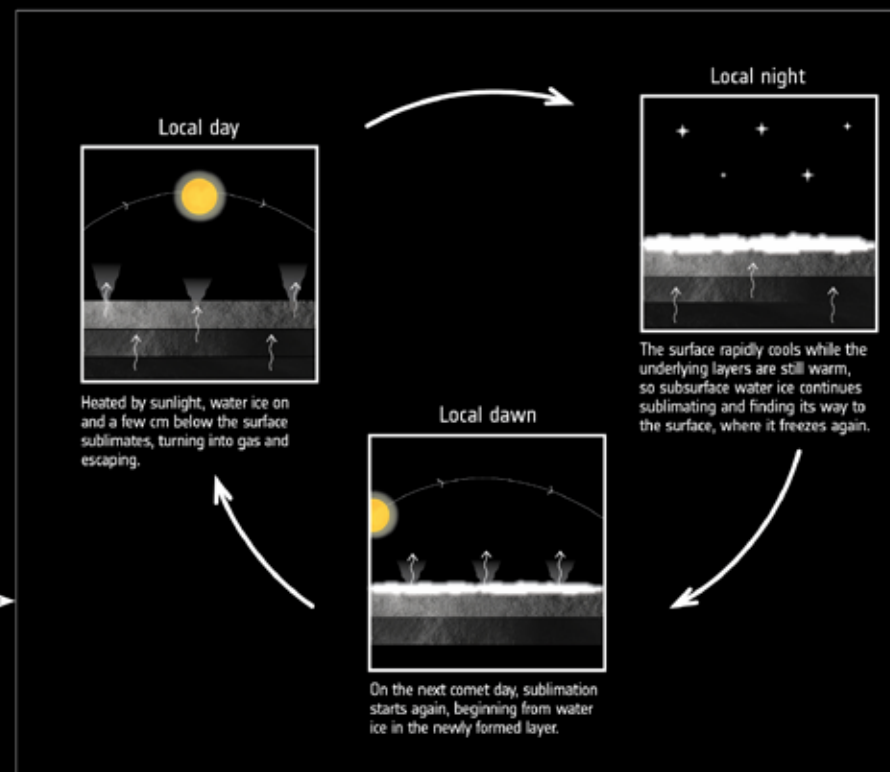
VIRTIS images

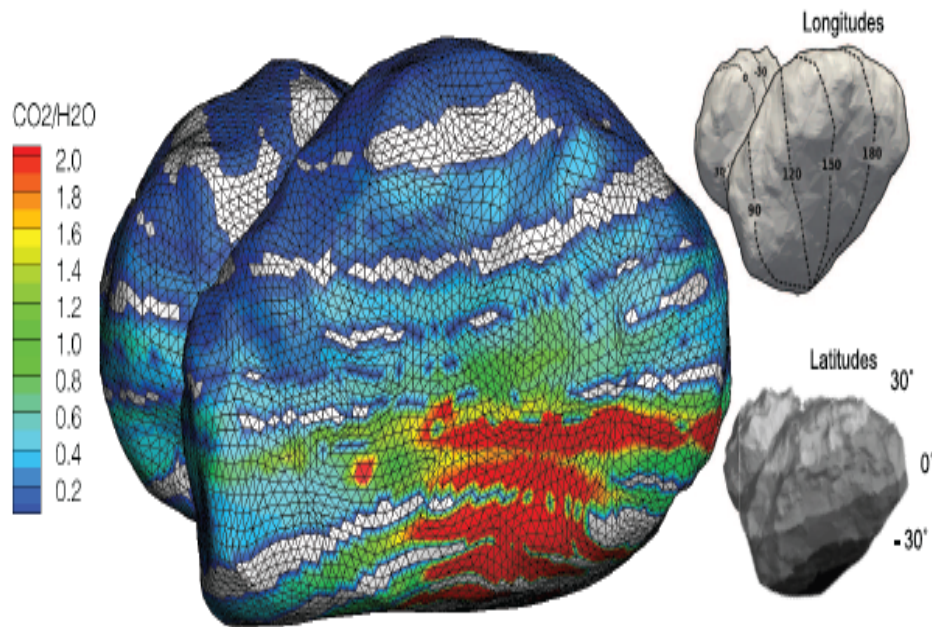
Ice abundance map

Temperature map

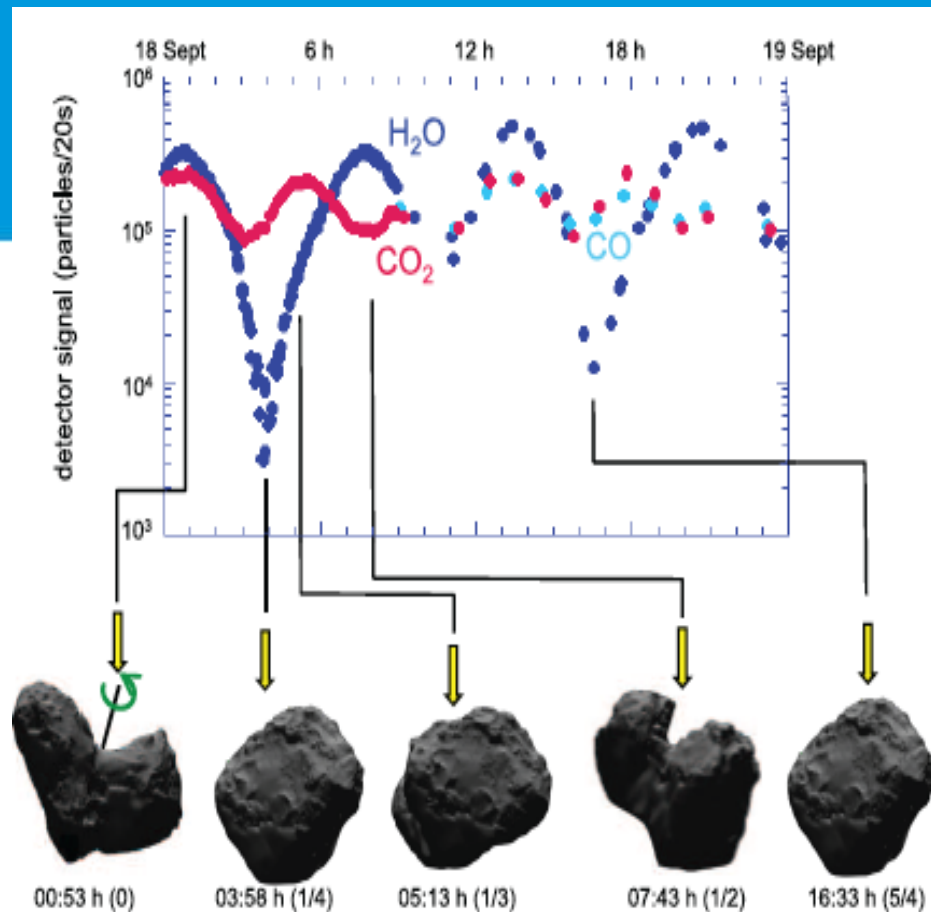


Water ice cycle at the comet





ESA/ROSETTA/ROSINA



Water (H₂O)

Carbon monoxide (CO)

Carbon dioxide (CO₂)

Ammonia (NH₃)

Methane (CH₄)

Methanol (CH₃OH)

Formaldehyde (CH₂O)

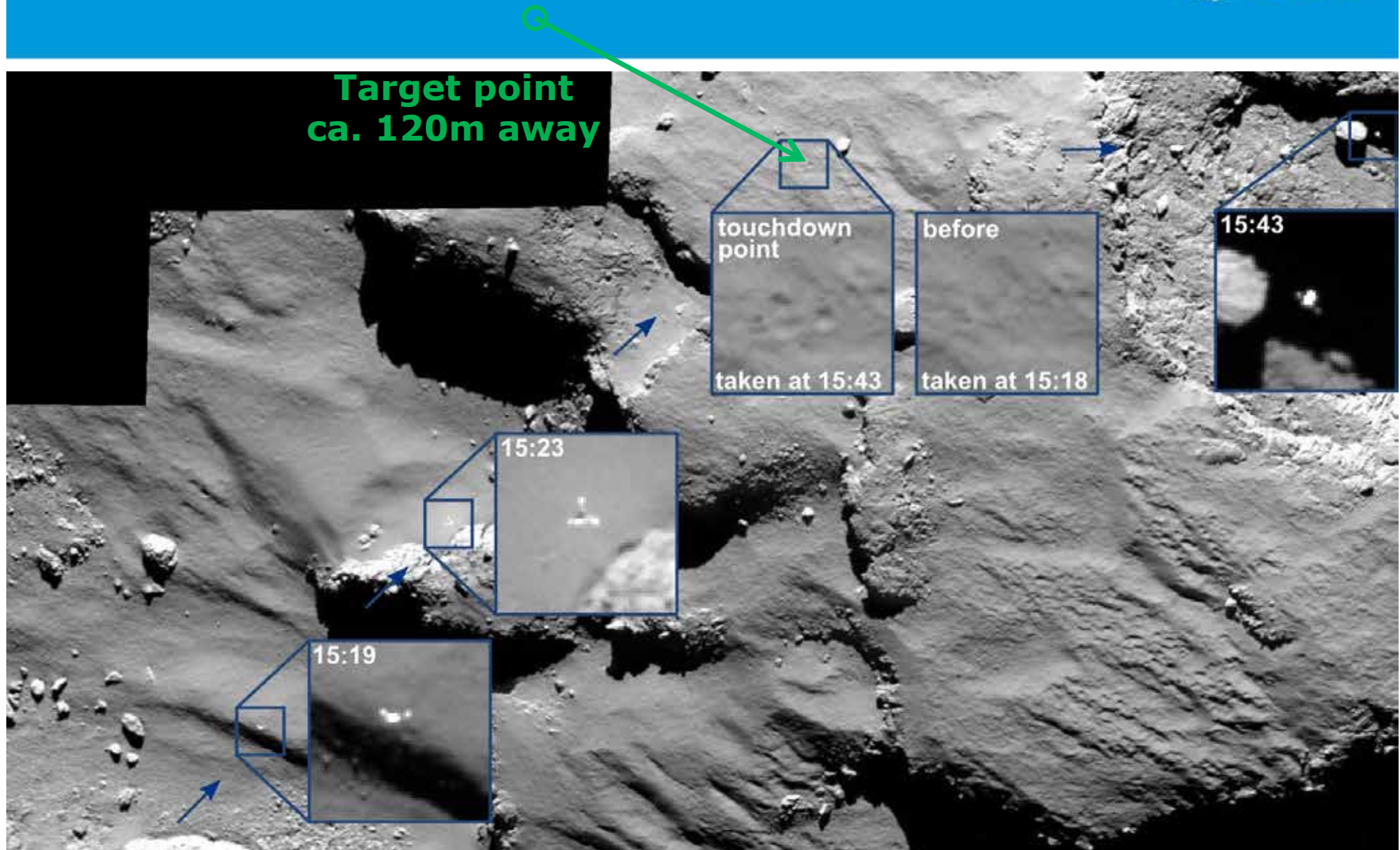
Hydrogen sulphide (H₂S)

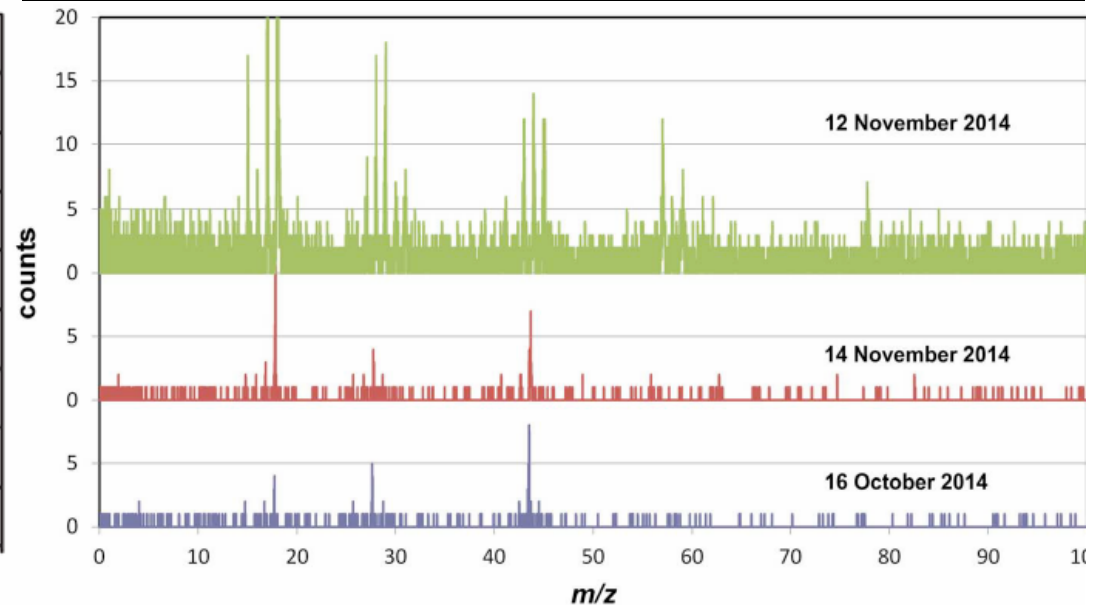
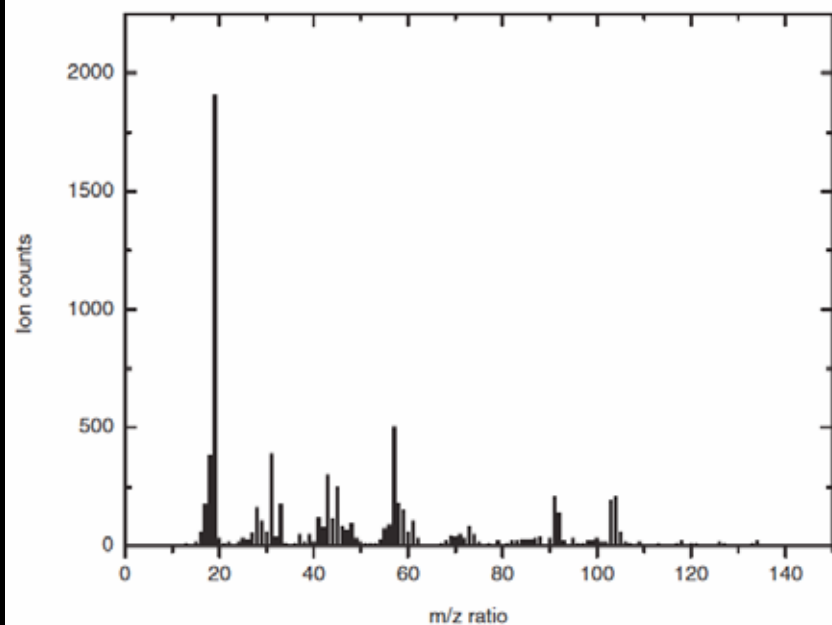
Hydrogen cyanide (HCN)

Sulphur dioxide (SO₂)

Carbon disulphide (CS₂)

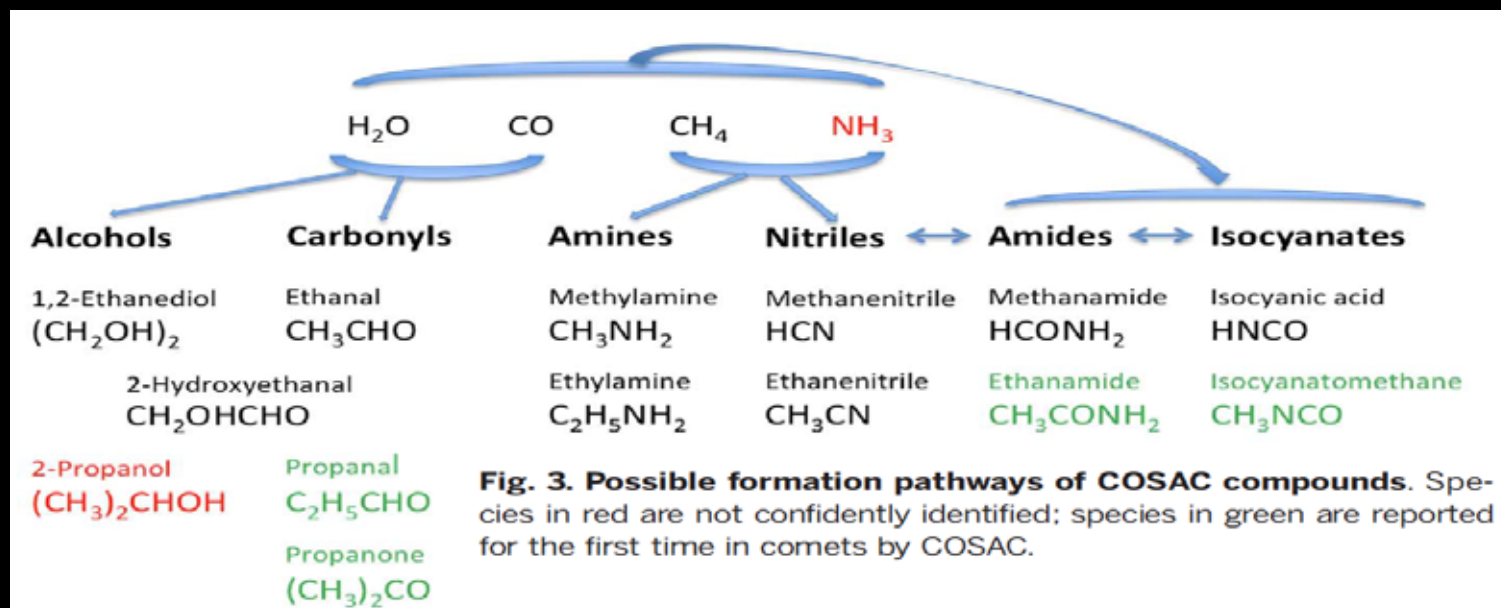
Landing





Wright et al., 2015

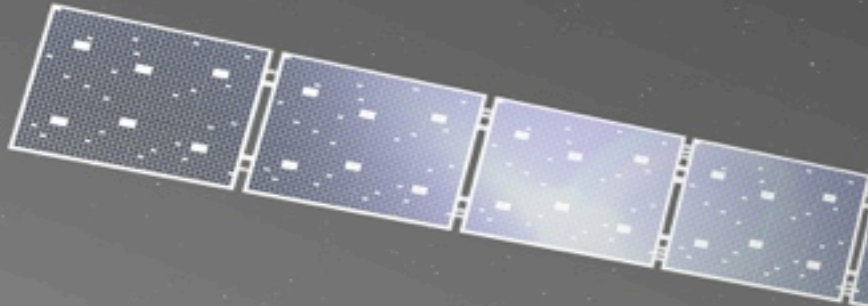
Goesmann et al., 2015



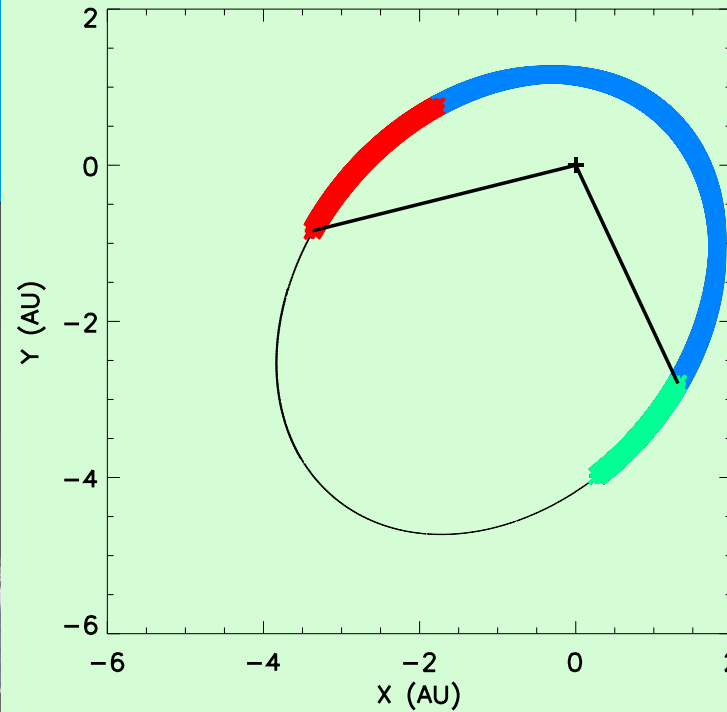


rosetta

<http://blogs.esa.int/rosetta/>



C67 P ORBIT J2000



January - August 2014

August 2014-December 2015

January - July 2016

Summer 2014

12 November 2014

August 2015

~~December 31, 2015~~

September 2016

Arrival at comet

Lander deployment

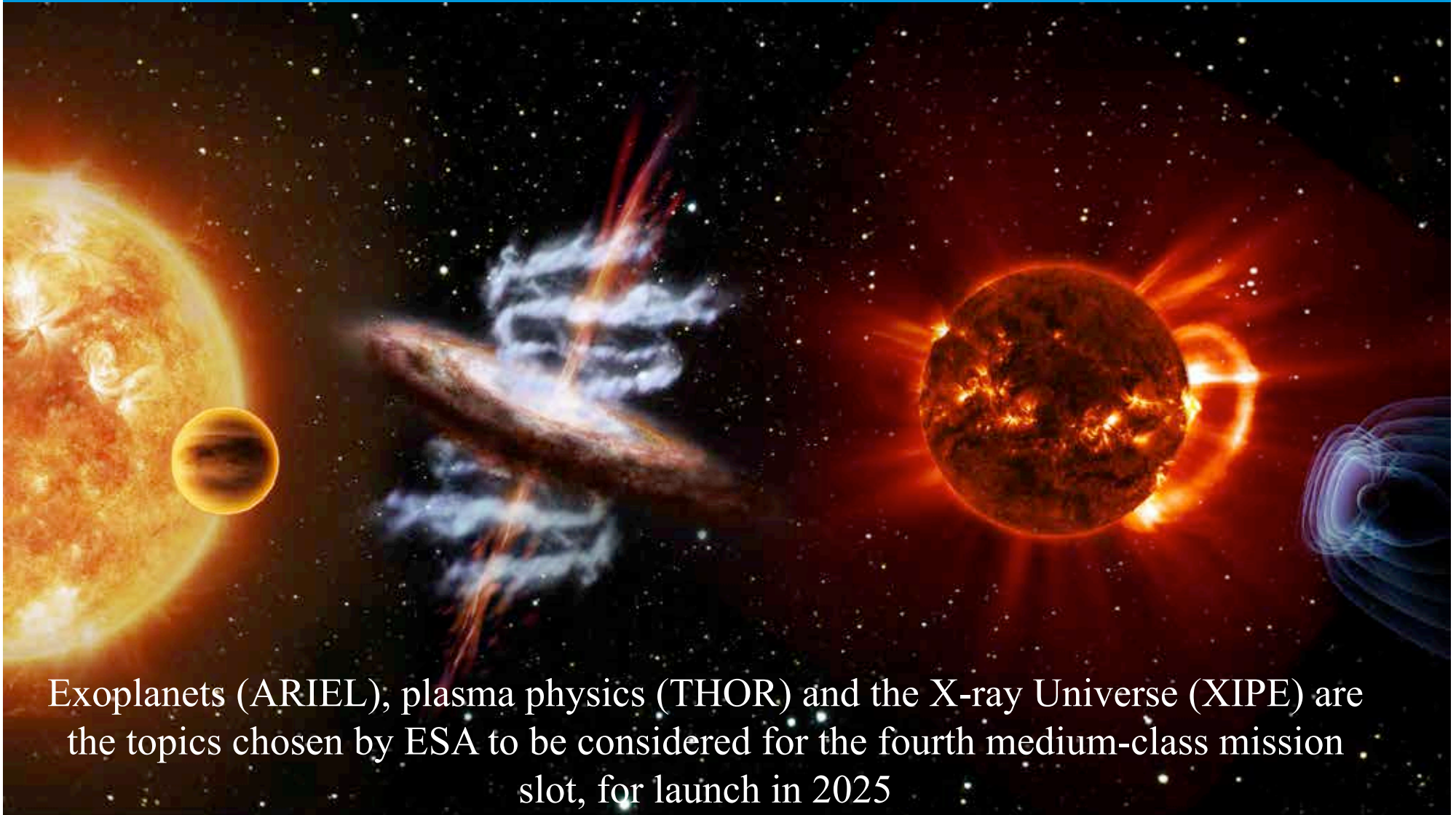
Perihelion

~~Nominal end of mission~~

End of extended mission

M4 call

Selection of missions for study



Exoplanets (ARIEL), plasma physics (THOR) and the X-ray Universe (XIPE) are the topics chosen by ESA to be considered for the fourth medium-class mission slot, for launch in 2025

M4 mission Present plan

A background image of a cosmic scene featuring a bright yellow-orange star on the left, a glowing spiral galaxy in the center, and a reddish-orange nebula on the right, all set against a dark space filled with distant stars.

Phase 0 (ESA internal CDF)	June to September 2015
Phase 0 completed	End September 2015
ITT for Phase A industrial studies	October 2015
Phase A kick-off	March 2016
Mission Selection Review completed (ARIEL, THOR, XIPE)	April 2017
Selection of M4 mission (SPC)	June 2017
Phase B1 kick-off of the selected M4 mission	July 2017
Phase B1 completed	September 2018
Mission Adoption Review	September/October 2018
Adoption of the M4 mission (SPC)	November 2018
Phase B2/C/D kick-off	July 2019
Launch	2026

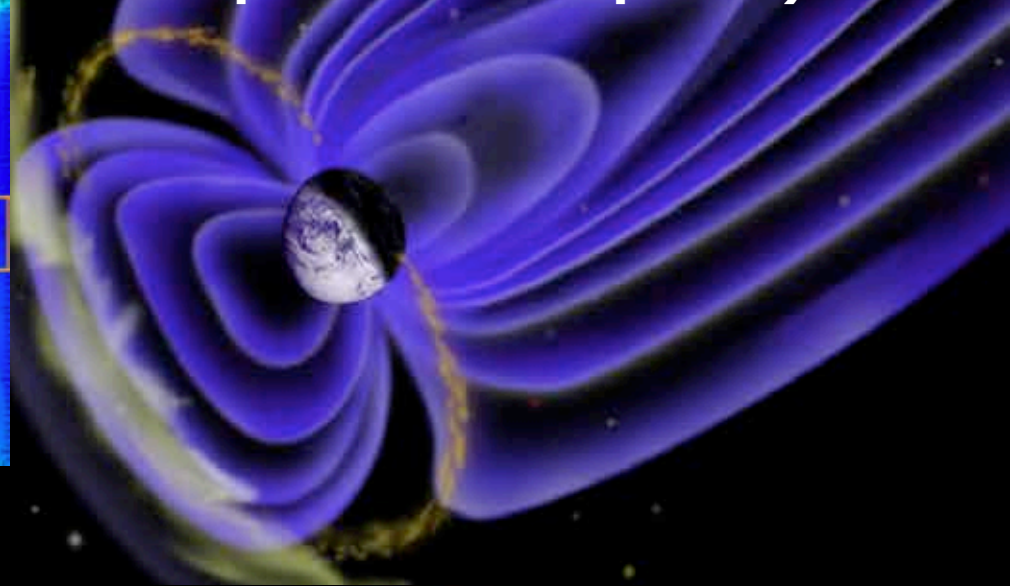
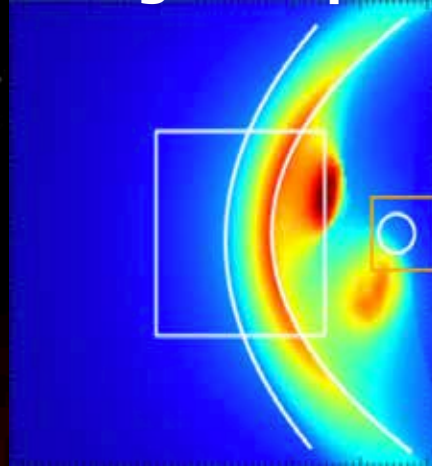


ESA – Chinese Academy of Sciences Joint mission



SMILE

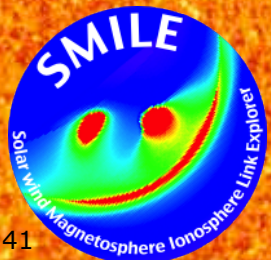
(Solar wind Magnetosphere Ionosphere Link Explorer)



Co-Pis: G. Branduardi-Raymont and C. Wang

Smile will investigate the interaction between Earth's protective shield – the magnetosphere – and the supersonic solar wind

Goal: understanding the physical processes taking place during the continuous interaction between the solar wind and the magnetosphere



- Planning of mission calls:

a.	M1, M2, L1	2007, 2007, 2007	slice 1
b.	M3, M4, L2	2010, 2014, 2014	slice 2
c.	M5, M6, L3	2015, 2017, 2017	slice 3
d.	M7	2020	...

All missions adopted during the first decade, leaving room for the preparation of the future beyond Cosmic Vision 10 years before.

- Planning of mission launches:

a.	M1, M2, L1	2018, 2020, 2022
b.	M3, M4, L2	2024, 2026, 2028
c.	M5, M6, L3	2030, 2032, 2034
d.	M7	2035

Missions

- [Show All Missions](#)

Cosmic Vision 2015–2025

- [Cosmic Vision](#)
- [Candidate Missions](#)
- [M-class Timeline](#)
- [L-class Timeline](#)

Cosmic Vision themes

- [Planets and Life](#)
- [The Solar System](#)
- [Fundamental Laws](#)
- [The Universe](#)
- [The Hot and Energetic Universe](#)

S-class mission

- [CHEOPS \[S1\]](#)



ANNOUNCEMENT OF THE PLANS FOR THE ISSUING OF A CALL FOR A MEDIUM-SIZE MISSION FOR LAUNCH IN 2029-2030 (M5)

20 July 2015

The Director of Science and Robotic Exploration of the European Space Agency plans to release, in late 2015 or early 2016, a Call for the M5 "Medium-size mission" with a planned launch date of 2029-2030 ("M5 Call" in the following).

The purpose of the present announcement is to inform the scientific community about the current planning, and to offer the scientific community the possibility of consulting the ESA Executive about their possible plans to submit proposals in response to the M5 Call.

The present announcement solicits non-binding Statements of Interest (SoI) from potential proposers. These Statements will allow the Executive to gauge the interest from the scientific community in the Call and to size the process accordingly. Submission of a SoI is not a pre-requisite for the eventual submission of a proposal in response to the actual M5 Call.

The content of the present announcement is non-binding, and does not commit ESA to release the M5 Call on any specific schedule. Also, the content and condition of the M5 Call, when finally released, may vary from the approach currently planned and described here. Thus, any information given here has to be considered as tentative and potentially subject to evolution, also as a function of the interaction between the Executive and the scientific community following the present announcement.



22-Oct-2015 07:21 UT

Shortcut URL

<http://sci.esa.int/jump.cf?m?oid=56198>

- The annual budget over five years is decided by unanimity at ESA Council at Ministerial level.
- Last full Ministerial was in 2012 in Naples.
- The ESA Council at Ministerial level, in Luxembourg in 2014, did not include the Science Programme in the Agenda.
- Next Ministerial meeting dealing with the Science Programme is planned before the end of 2016 (in Switzerland).
- The next Council at Ministerial level is an important opportunity, the first to take place during the implementation time of Cosmic Vision.